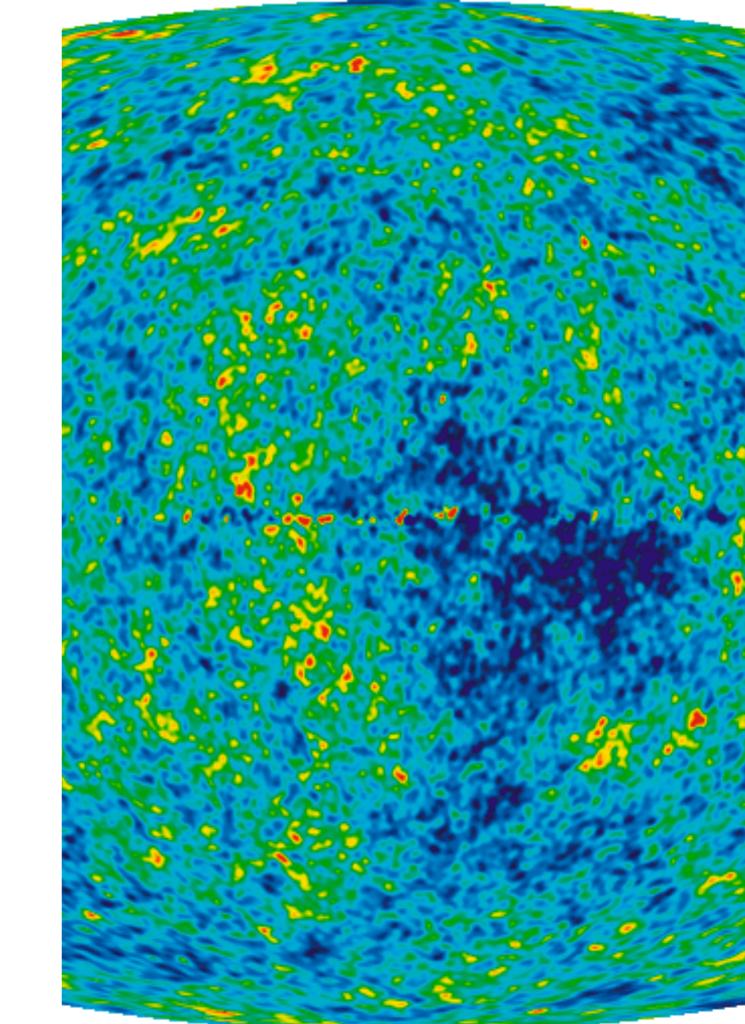
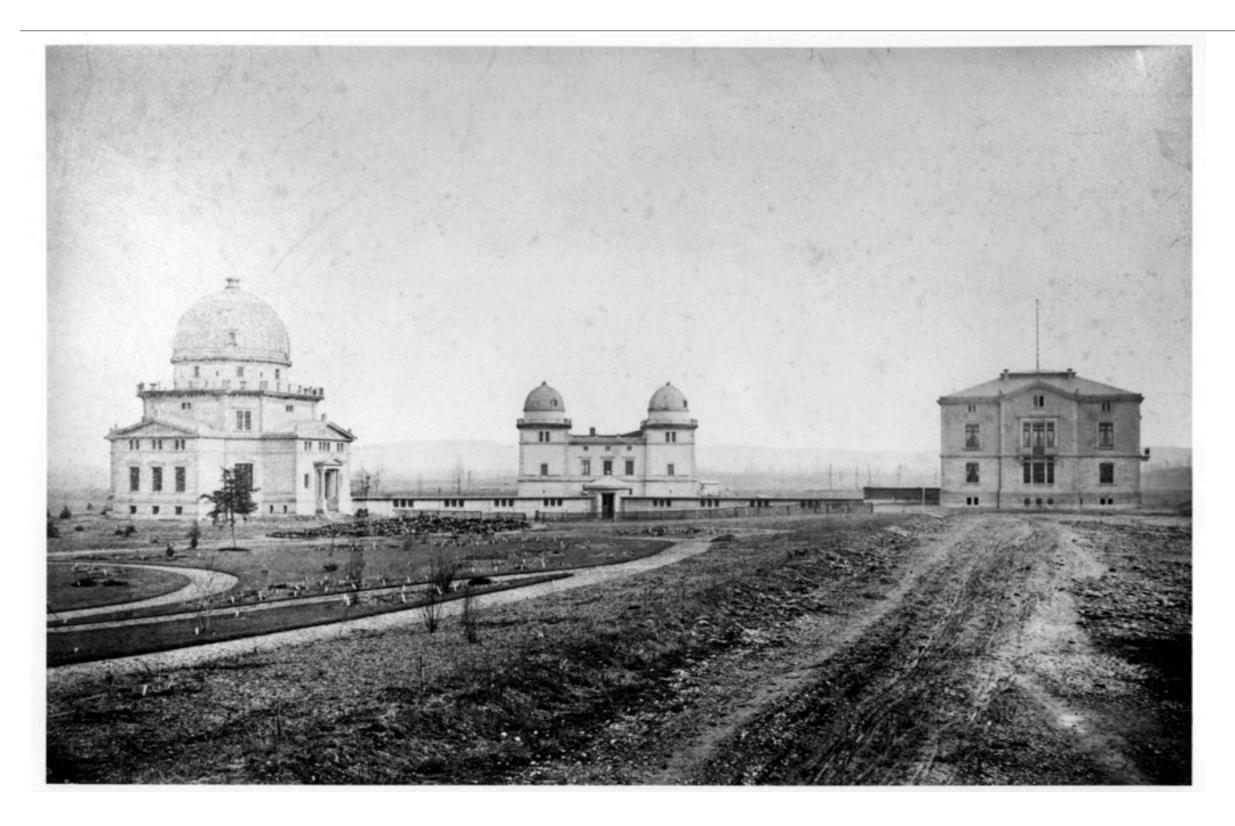
The Cosmological Sky

Dominique Aubert

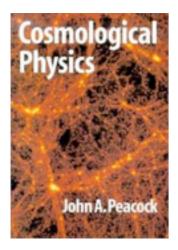


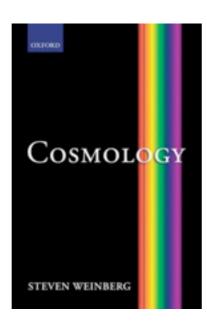
Welcome to the Observatory

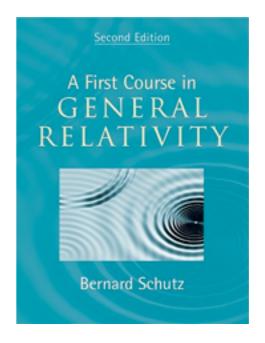


Bibliography

•	Ned Wright website : <u>http://www.astro.ucla.edu/~wright/cosmolog.htm</u>
•	Wayne Hu website : <u>http://background.uchicago.edu/~whu/</u>
•	WMAP Papers : <u>http://map.gsfc.nasa.gov/news/tp_links.html</u>
•	IPAC level 5: <u>http://ned.ipac.caltech.edu/level5/</u>
•	Cosmological Physics by J. Peacock
•	Cosmology by S. Weinberg
•	A First Course in General Relativity By Bernard Schutz

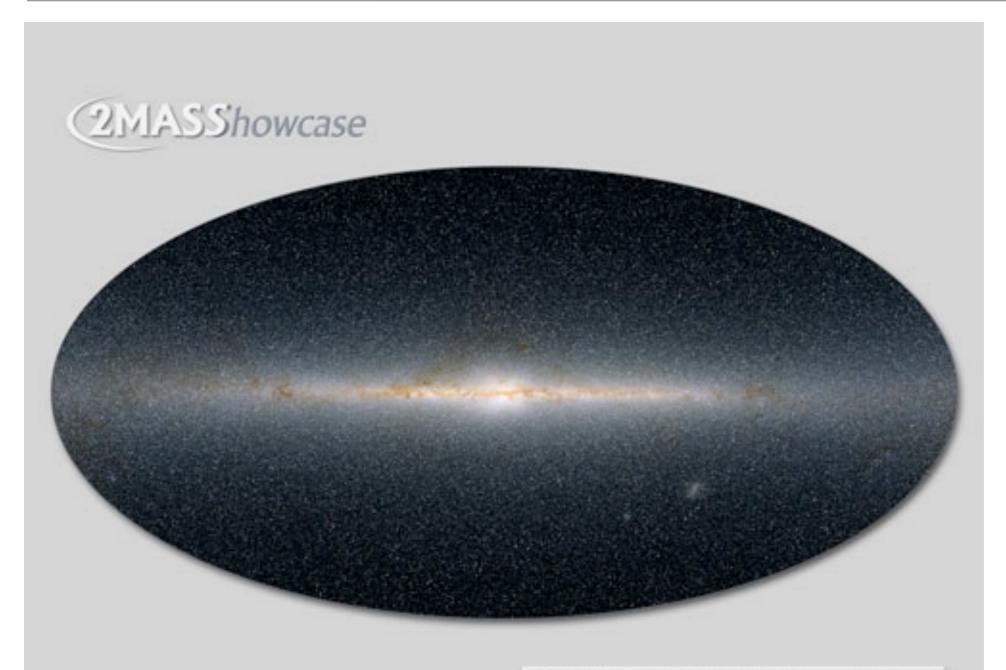






Zooming Out

The Galaxy

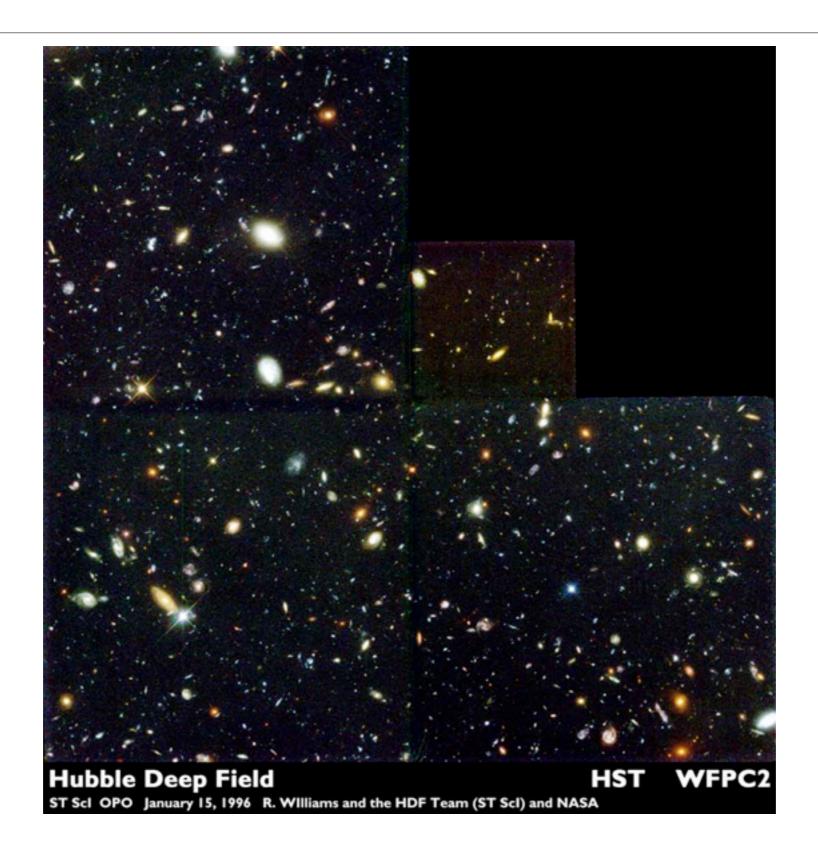


The Infrared Milky Way This map of the

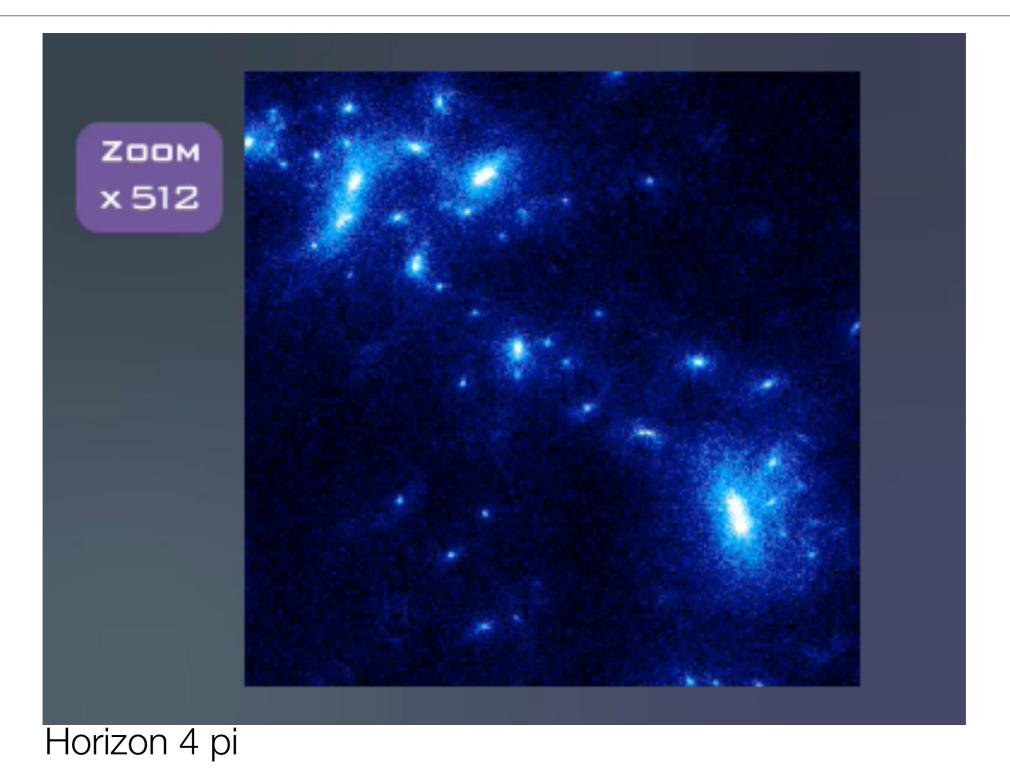
infrared sky includes the light of a half billion stars

Two Micron All Sky Survey Image Mosaic: Infrared Processing and Analysis Center/Caltech & University of Massachusetts

HDF

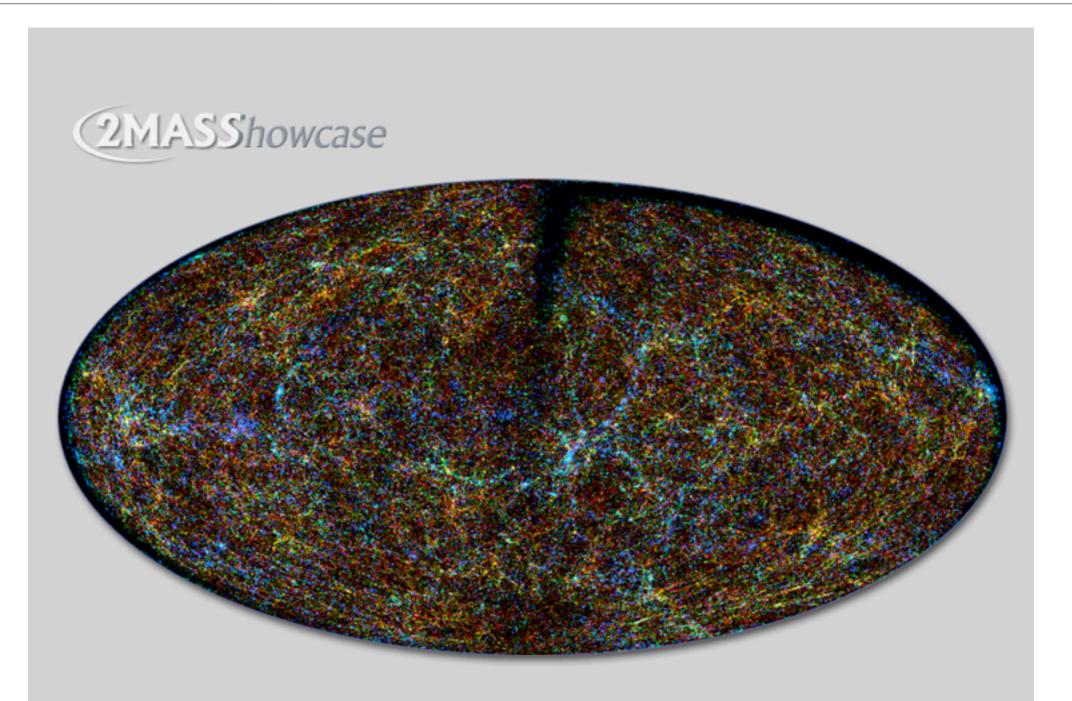


Large Scale Structures : LSS



Credit : Horizon Project

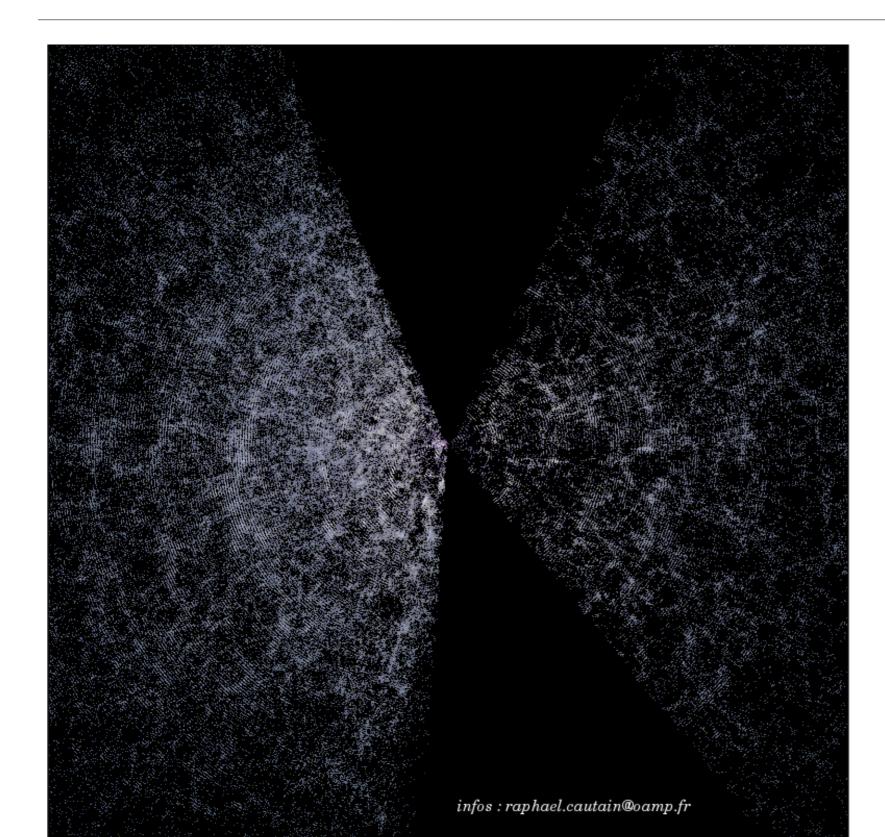
LSS



Galaxies of the Infrared Sky Near and far structures in the local universe are color-coded by galaxy brightness

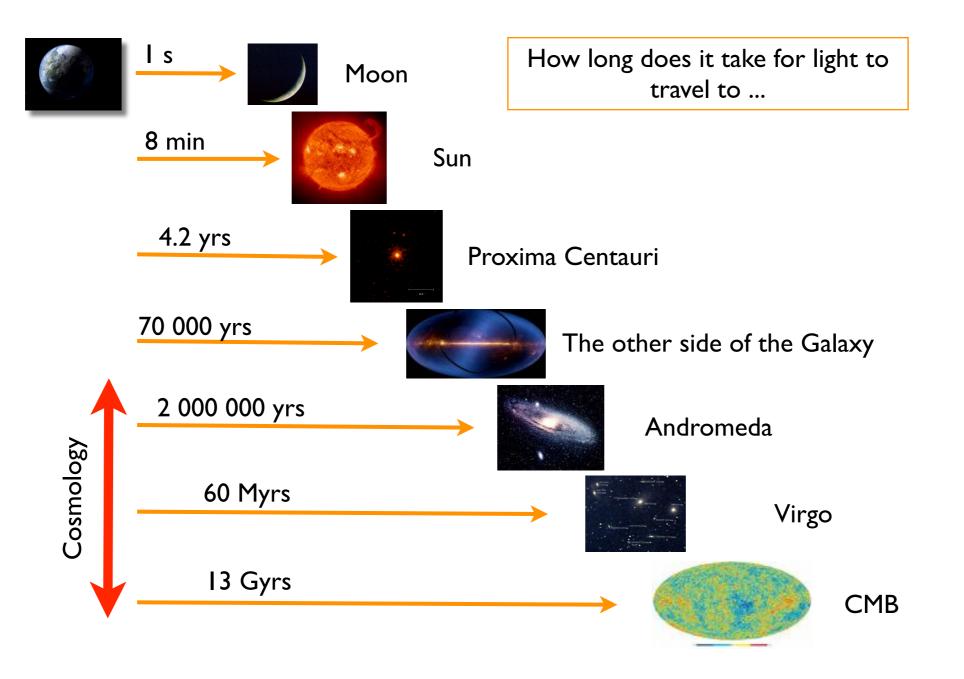
Two Micron All Sky Survey Image Mosaic: Infrared Processing and Analysis Center/Caltech & University of Massachusetts

LSS

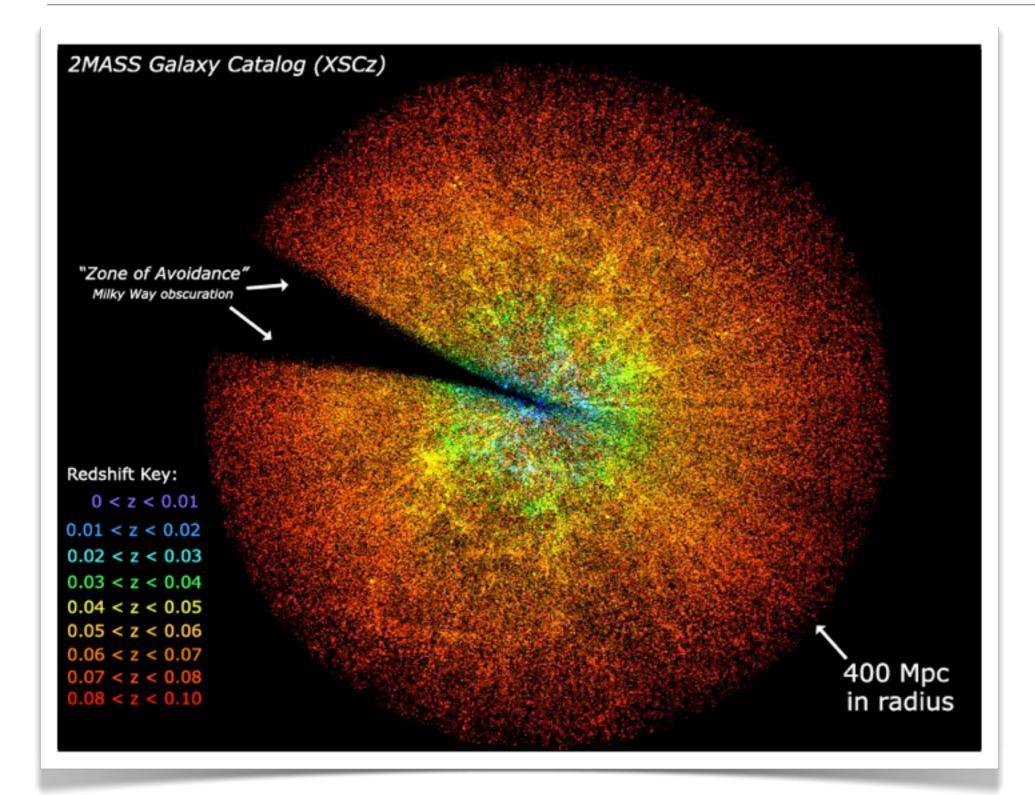


Voids... Filaments... Clusters...

Cosmological distances



Universe Expansion



the reddening of light or «redshift» is related to the motion of emitters

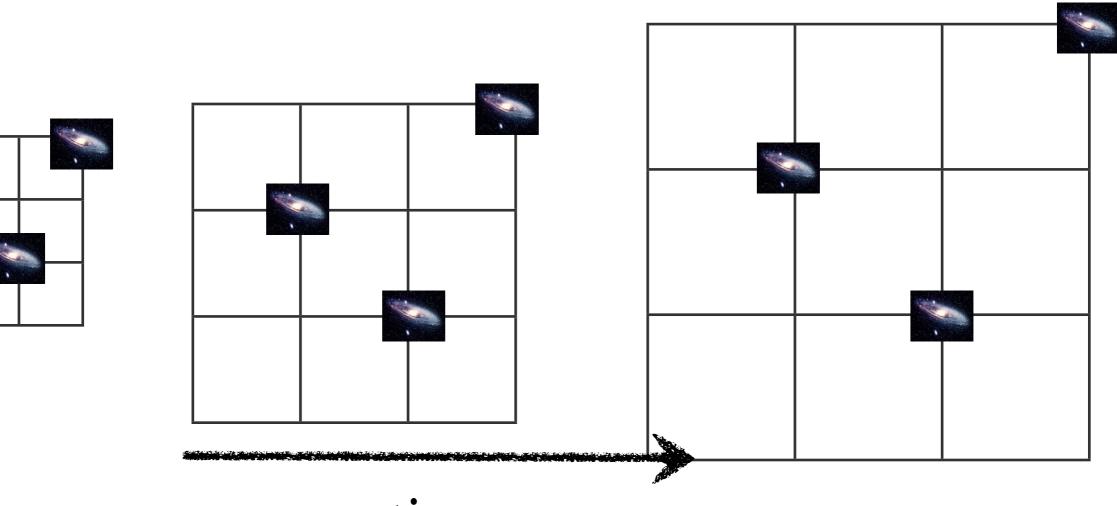
Cosmological Principle

The <u>standard</u> cosmological model assumes :

the Universe is isotropic and homogeneous (on cosmological scales)
Gravitation is described by General Relativity (GR)

Extension of the Copernician principle. It leads to:

Expansion



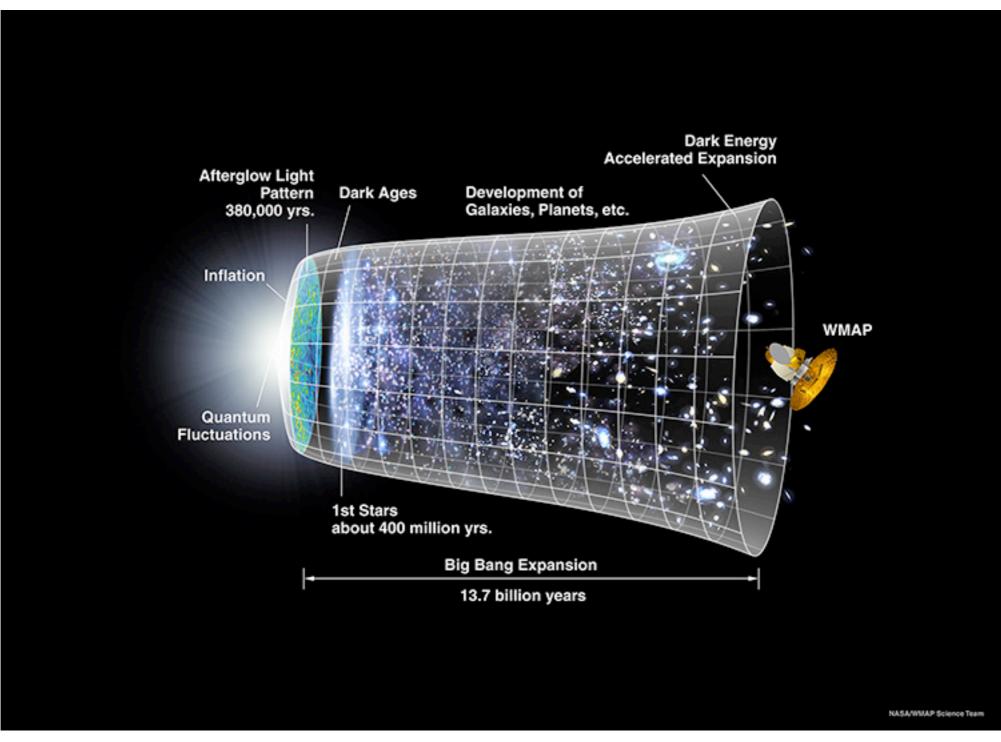
time

as we will discuss later:

- space expands with time

- the Universe temperature decreases with time

History of the Universe



Nasa /WMAP team

The standard model

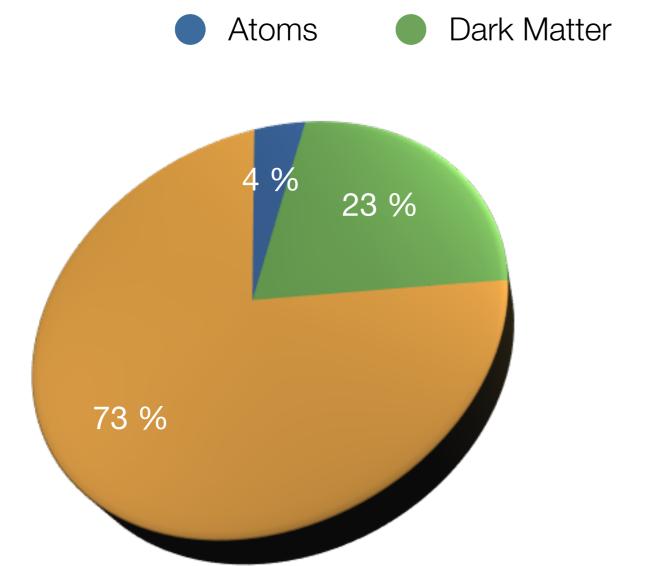
THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 192:18 (47pp), 2011 February

KOMATSU ET AL.

Class	Parameter	WMAP Seven-year MLb	WMAP+BAO+H0 ML	WMAP Seven-year Mean ^c	WMAP+BAO+H0 Mean
Primary	$100\Omega_b h^2$	2.227	2.253	2.249+0.056	2.255 ± 0.054
	$\Omega_c h^2$	0.1116	0.1122	0.1120 ± 0.0056	0.1126 ± 0.0036
	Ω_{Λ}	0.729	0.728	$0.727^{+0.030}_{-0.029}$	0.725 ± 0.016
	n_S	0.966	0.967	0.967 ± 0.014	0.968 ± 0.012
	τ	0.085	0.085	0.088 ± 0.015	0.088 ± 0.014
	$\Delta_R^2(k_0)^d$	2.42×10^{-9}	2.42×10^{-9}	$(2.43 \pm 0.11) \times 10^{-9}$	$(2.430 \pm 0.091) \times 10^{-9}$
Derived	σ_8	0.809	0.810	0.811+0.030 -0.031	0.816 ± 0.024
	H_0	70.3 km s ⁻¹ Mpc ⁻¹	70.4 km s ⁻¹ Mpc ⁻¹	$70.4 \pm 2.5 \text{ km s}^{-1} \text{ Mpc}^{-1}$	$70.2 \pm 1.4 \text{ km s}^{-1} \text{ Mpc}^{-1}$
	Ω_b	0.0451	0.0455	0.0455 ± 0.0028	0.0458 ± 0.0016
	Ω_c	0.226	0.226	0.228 ± 0.027	0.229 ± 0.015
	$\Omega_m h^2$	0.1338	0.1347	0.1345+0.0056 -0.0055	0.1352 ± 0.0036
	Zreion ^e	10.4	10.3	10.6 ± 1.2	10.6 ± 1.2
	tof	13.79 Gyr	13.76 Gyr	13.77 ± 0.13 Gyr	13.76 ± 0.11 Gyr
Amount of Baryonic matter			Amount of D	Dark matter	
		-			
«Ge	Geometrical» distribution of initial seeds			Amount of D	ark Energy
	II IIIIai	SEEUS	Instant of E	Reionisation	
			IIISTALL OF L	neiui IISaliui I	

Table 1 Summary of the Cosmological Parameters of ACDM Model^a

Energetic content of the Universe, today



 $\Omega_{\Lambda} = 0.73$ $\Omega_{m} = 0.27$ $\Omega_{b} = 0.04$ $\Omega_{r} \sim 0.00001$

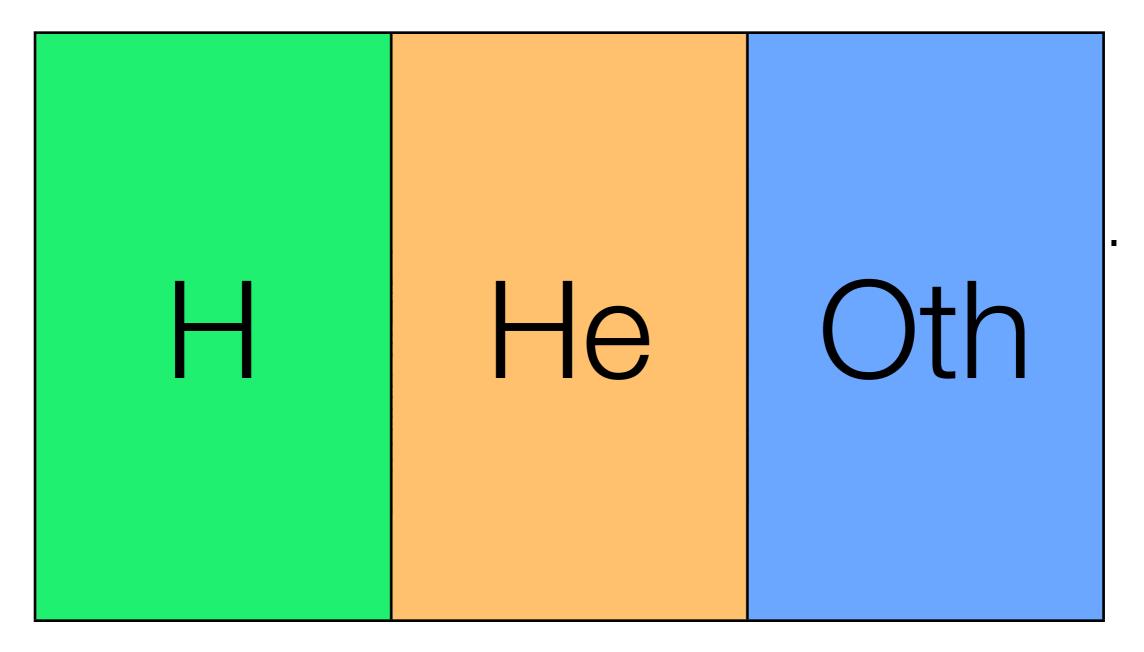
$$\Omega_m + \Omega_r + \Omega_\Lambda = 1$$

Vacuum Energy



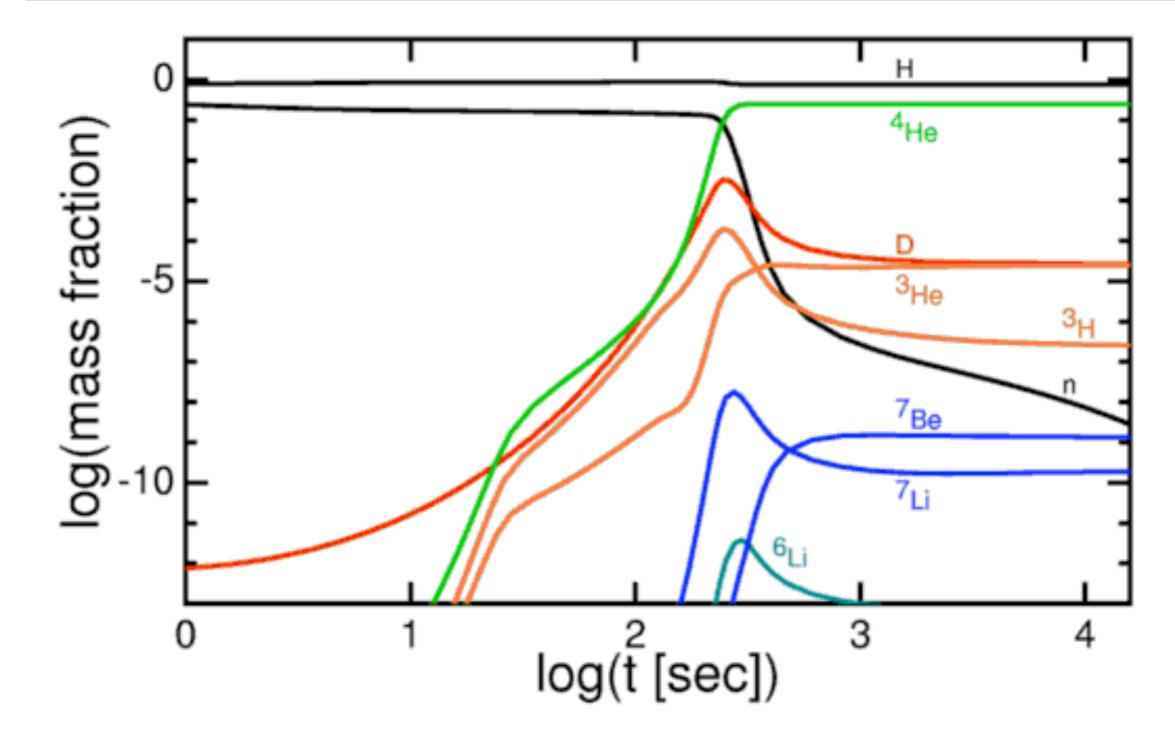
Baryons (or the story of the first 3 minutes)

Light Elements in the Universe



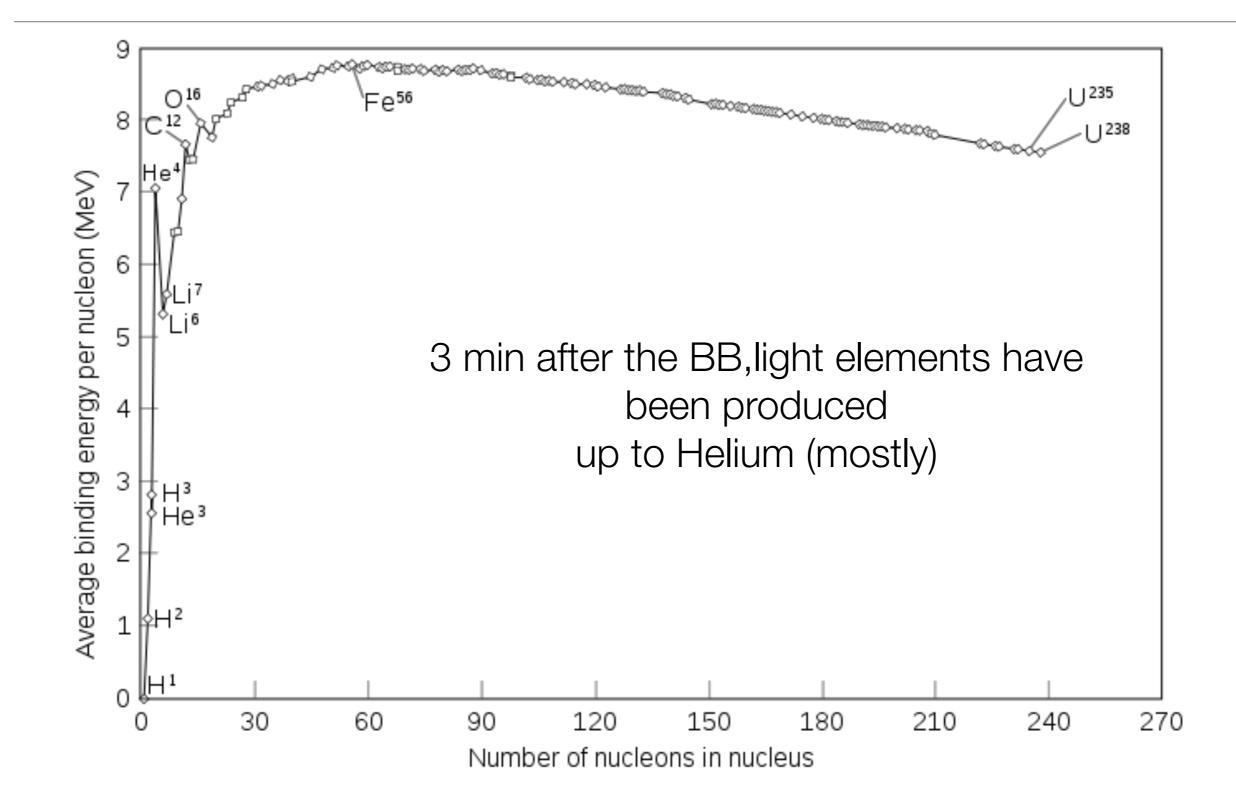
75 % H -- 25% He -- 0% Others (Metals)

BBN

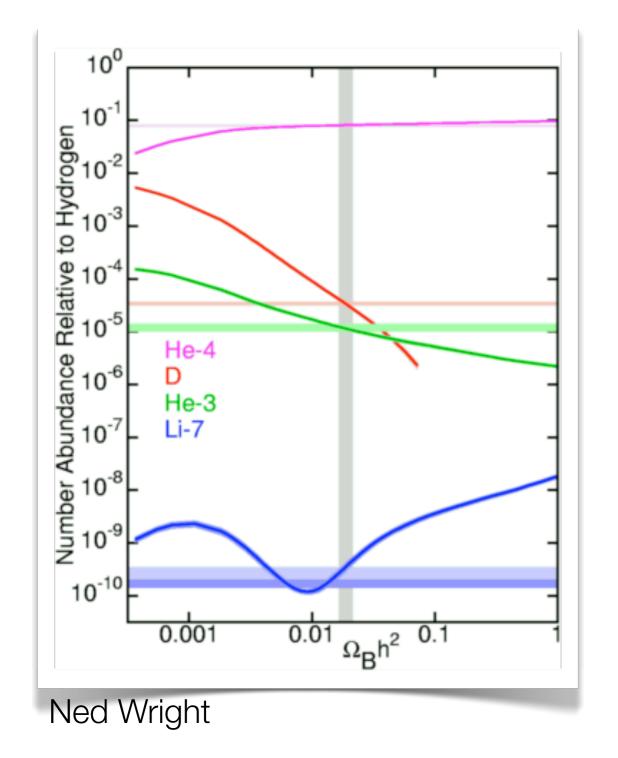


as the Universe cools down Deuterium than Helium can be created

Stability of Nuclei



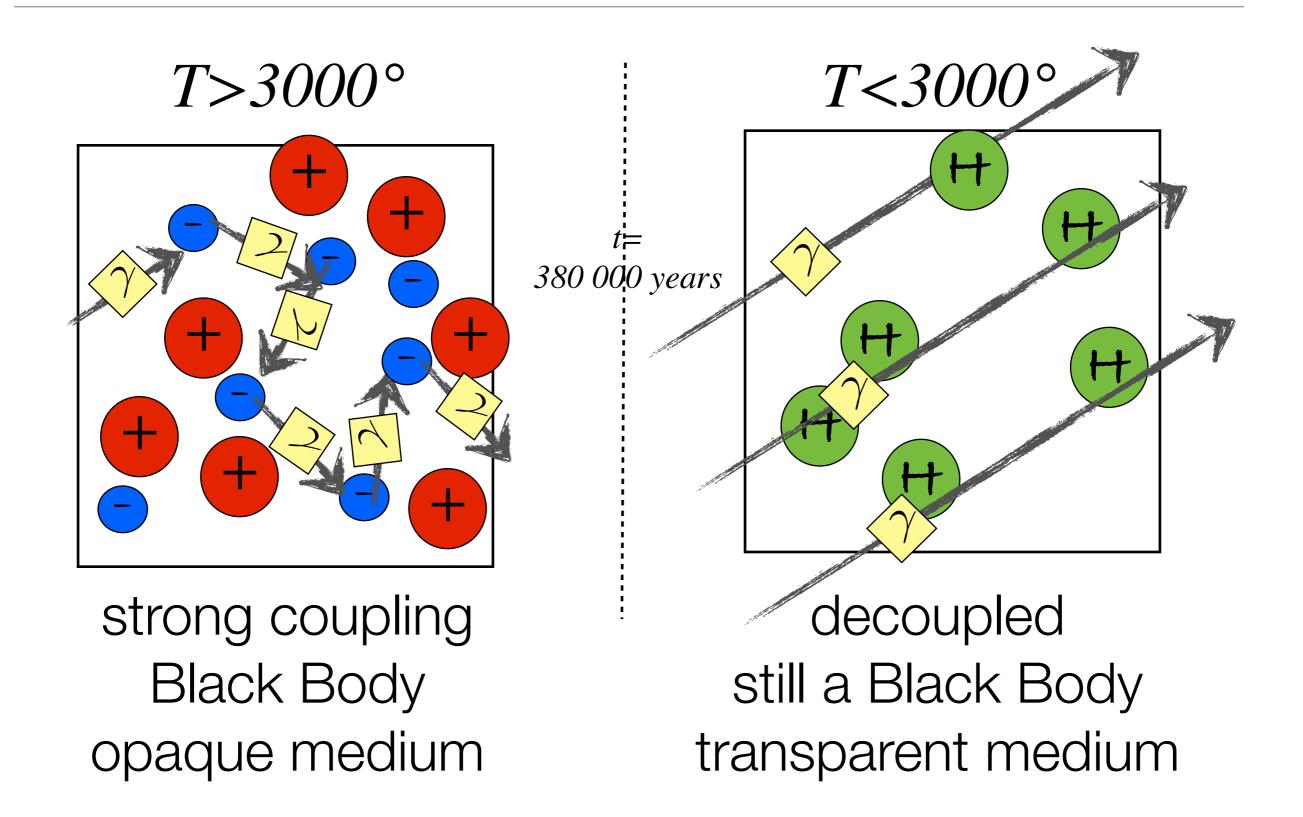
Big Bang Nucleosynthesis



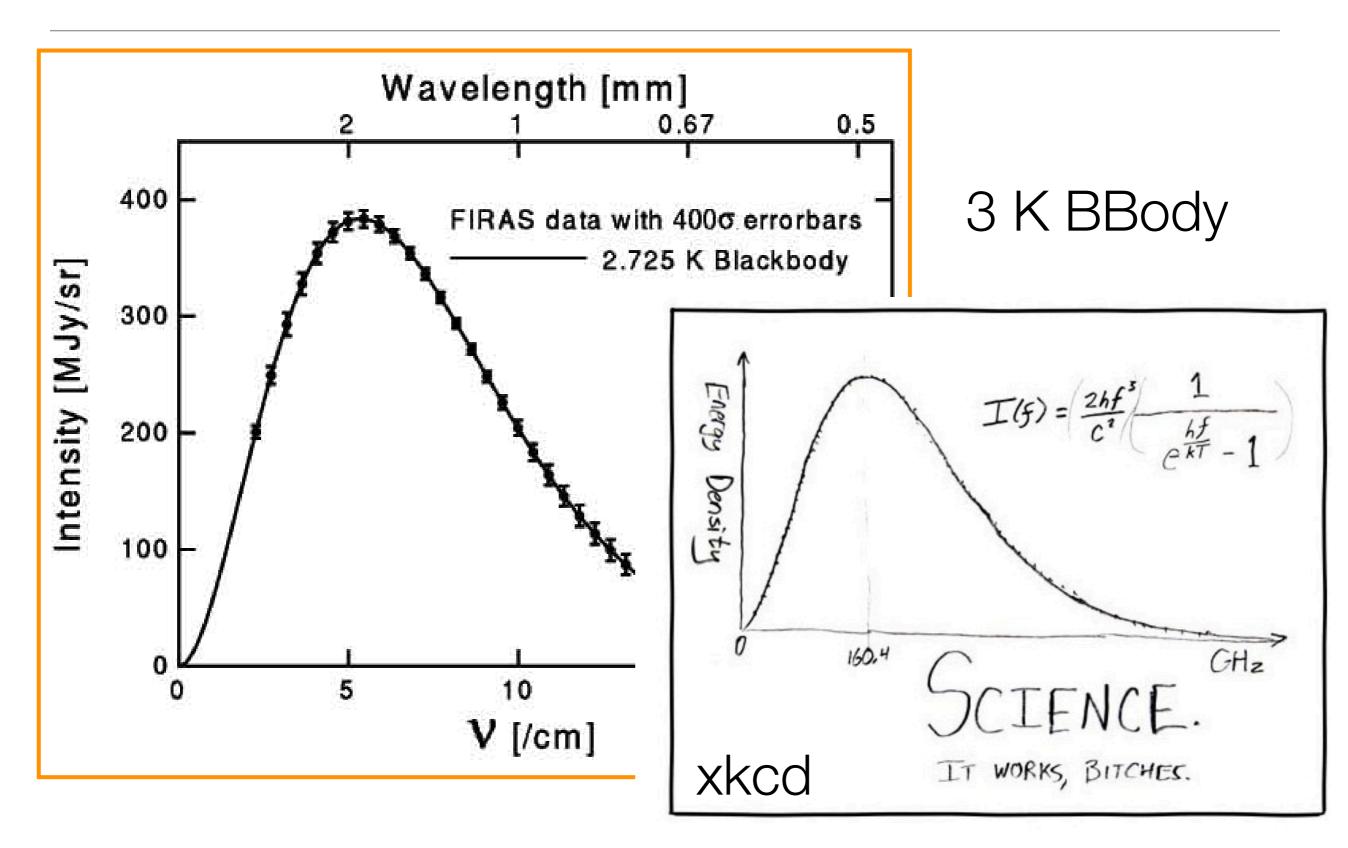
The total amount of Baryons is constrained by BBN

Light (or the first 400 000 years)

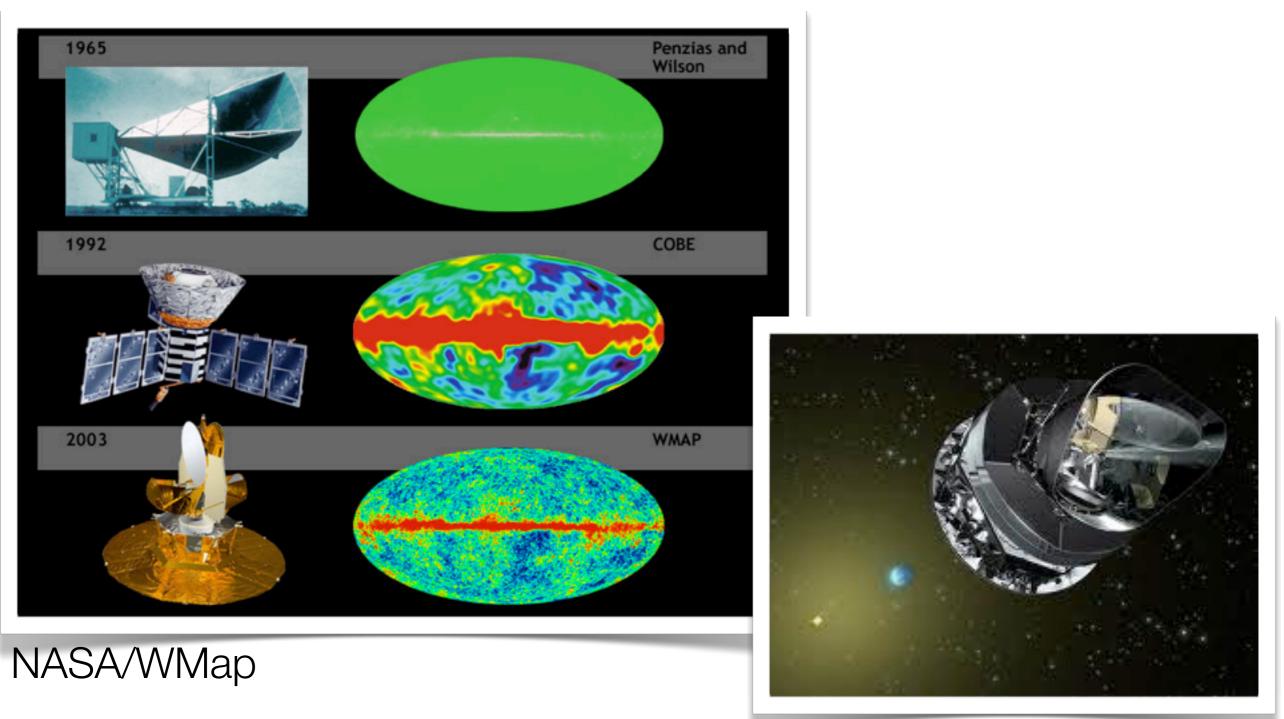
Last Scattering



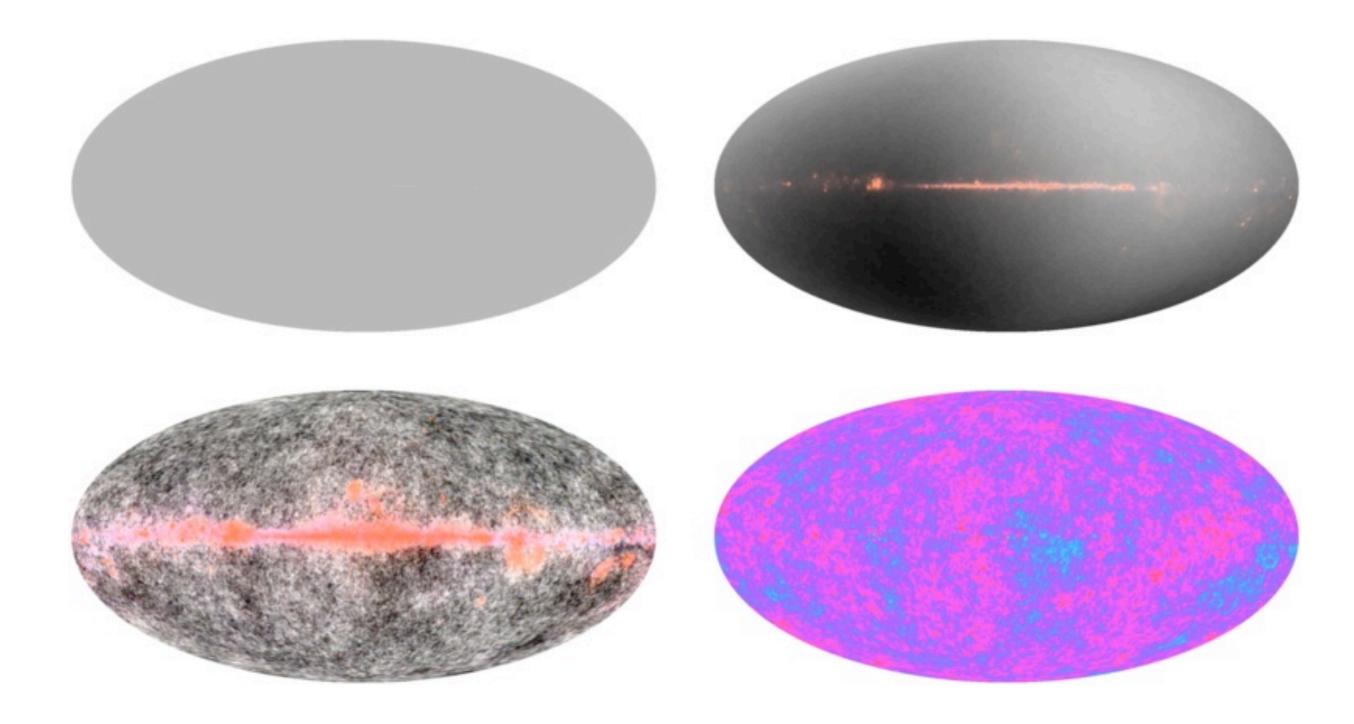
The CMB : the perfect Black Body



Looking for the CMB

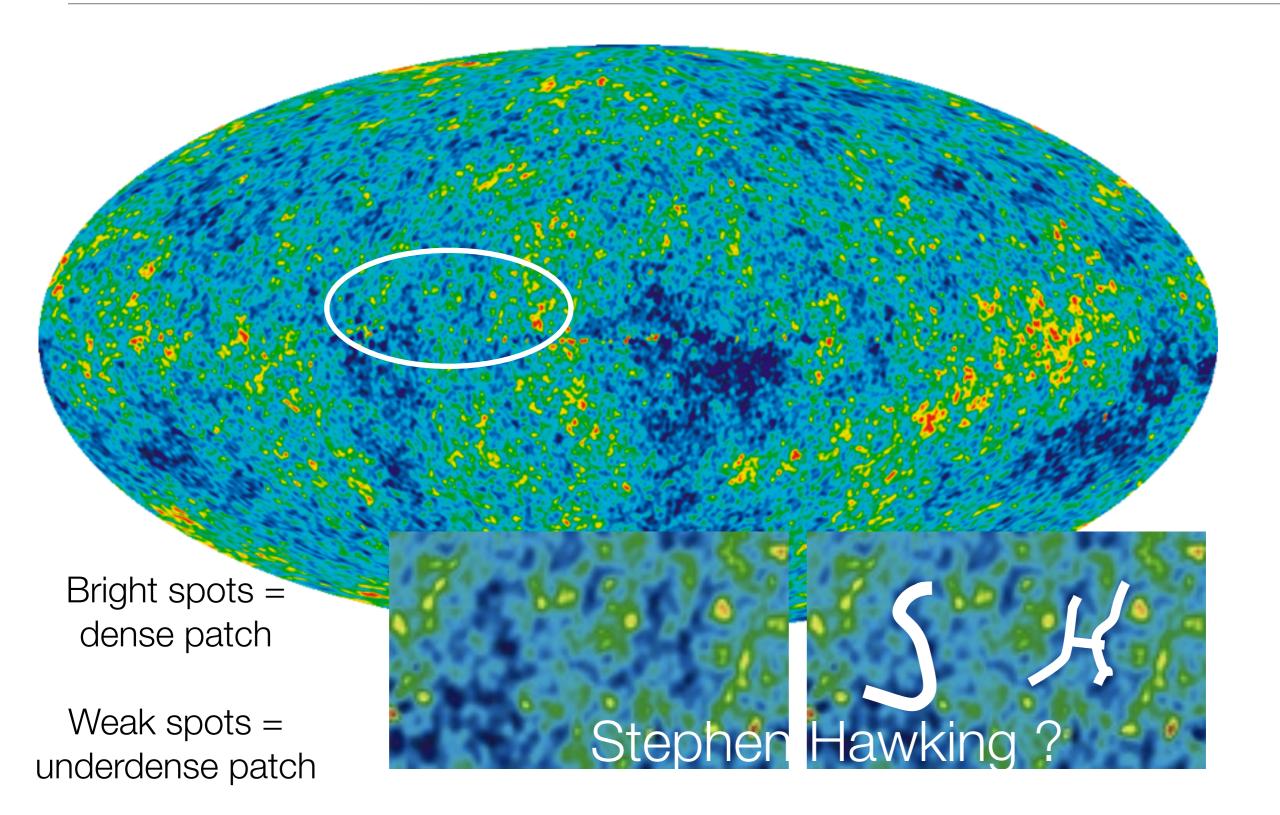




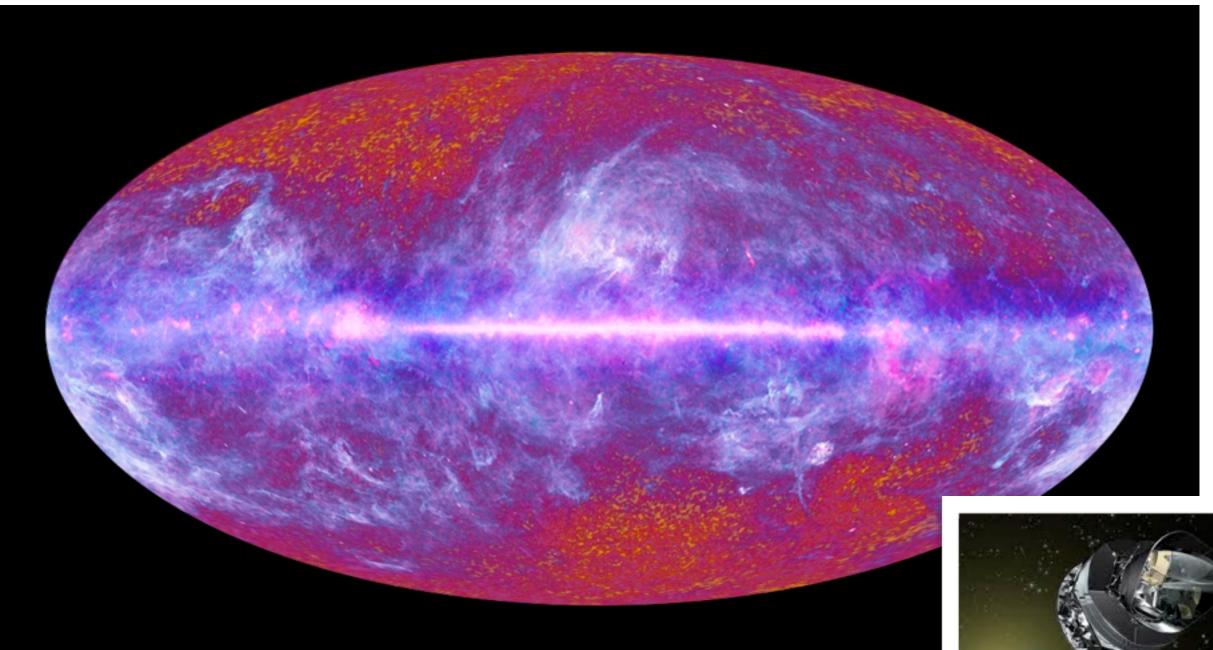


http://www.astro.ucla.edu/~wright/

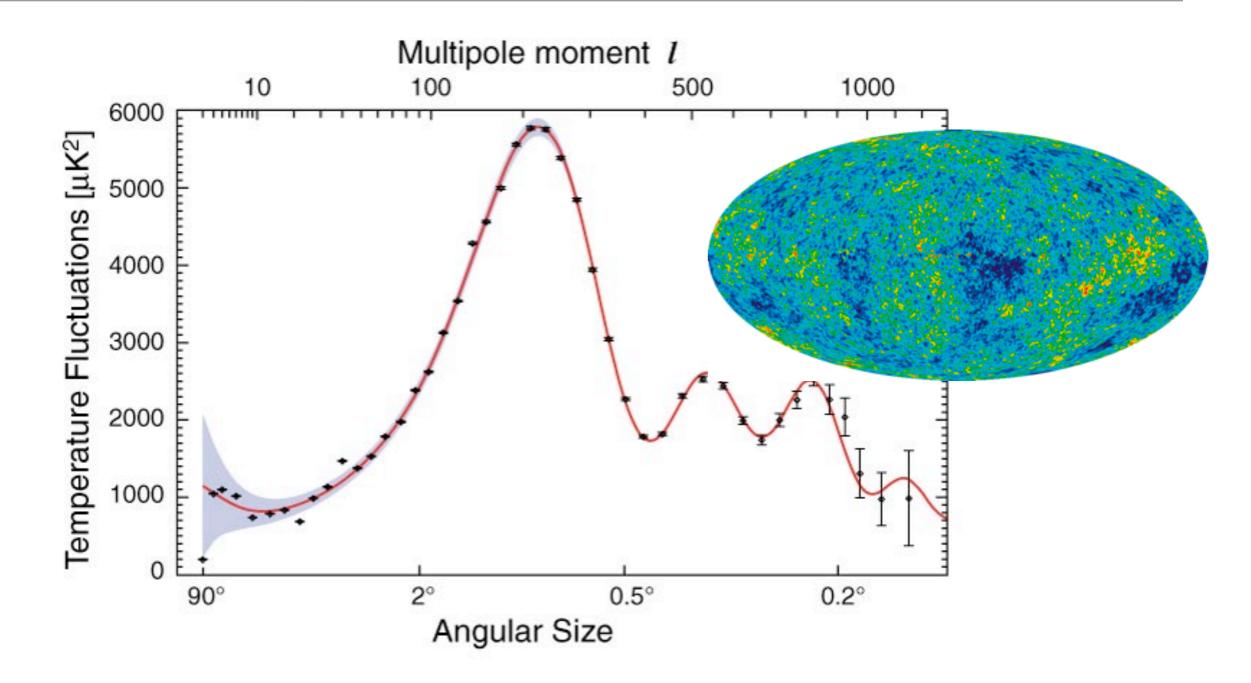
The CMB



Waiting for Planck ...

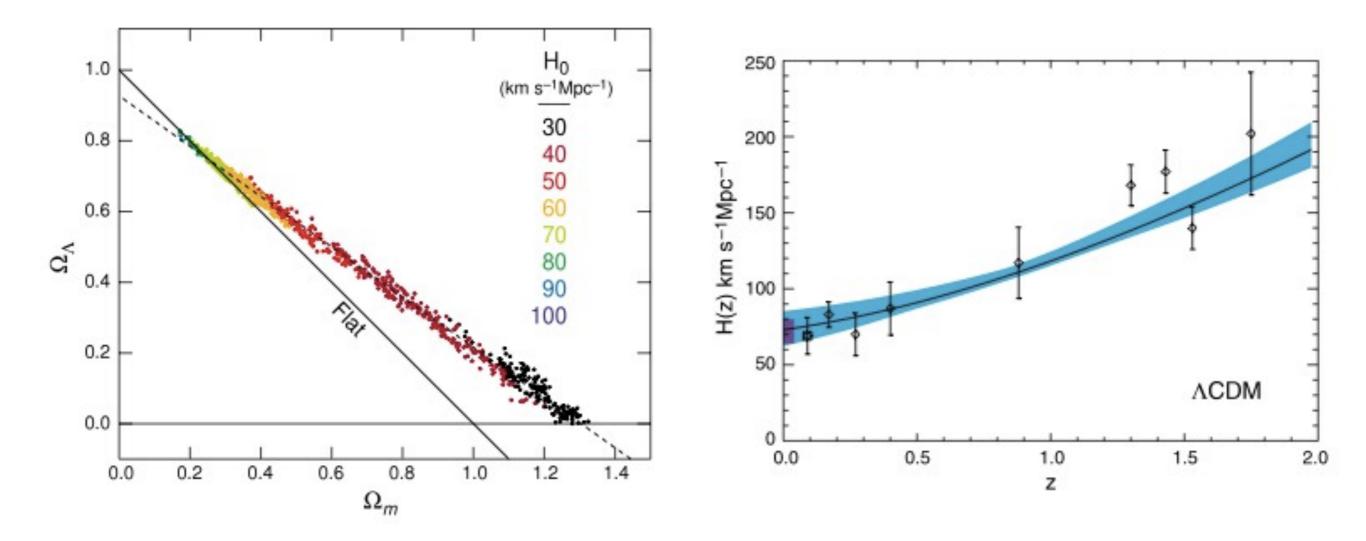


The CMB power spectrum



Describes the relative intensity of «spots» in CMB maps

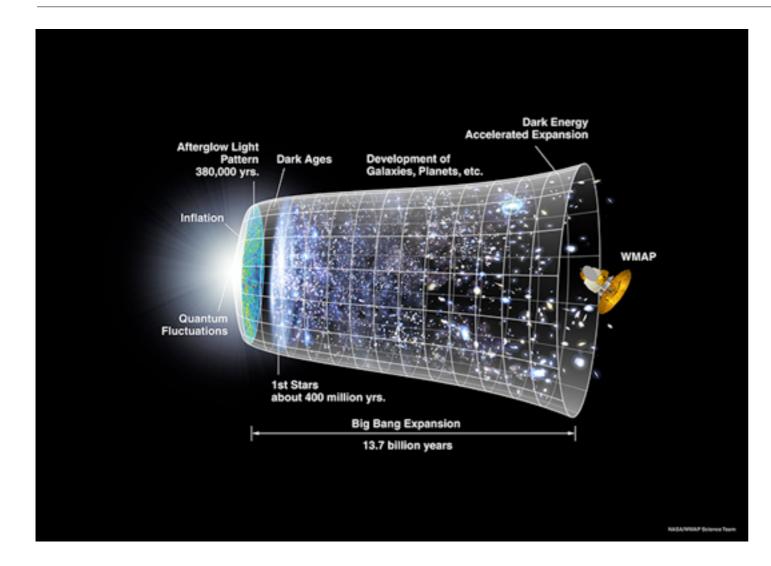
Flatness of the Universe



Spergel et al. 2007 (WMAP 3)

Stars & Reioniation

The Dark ages



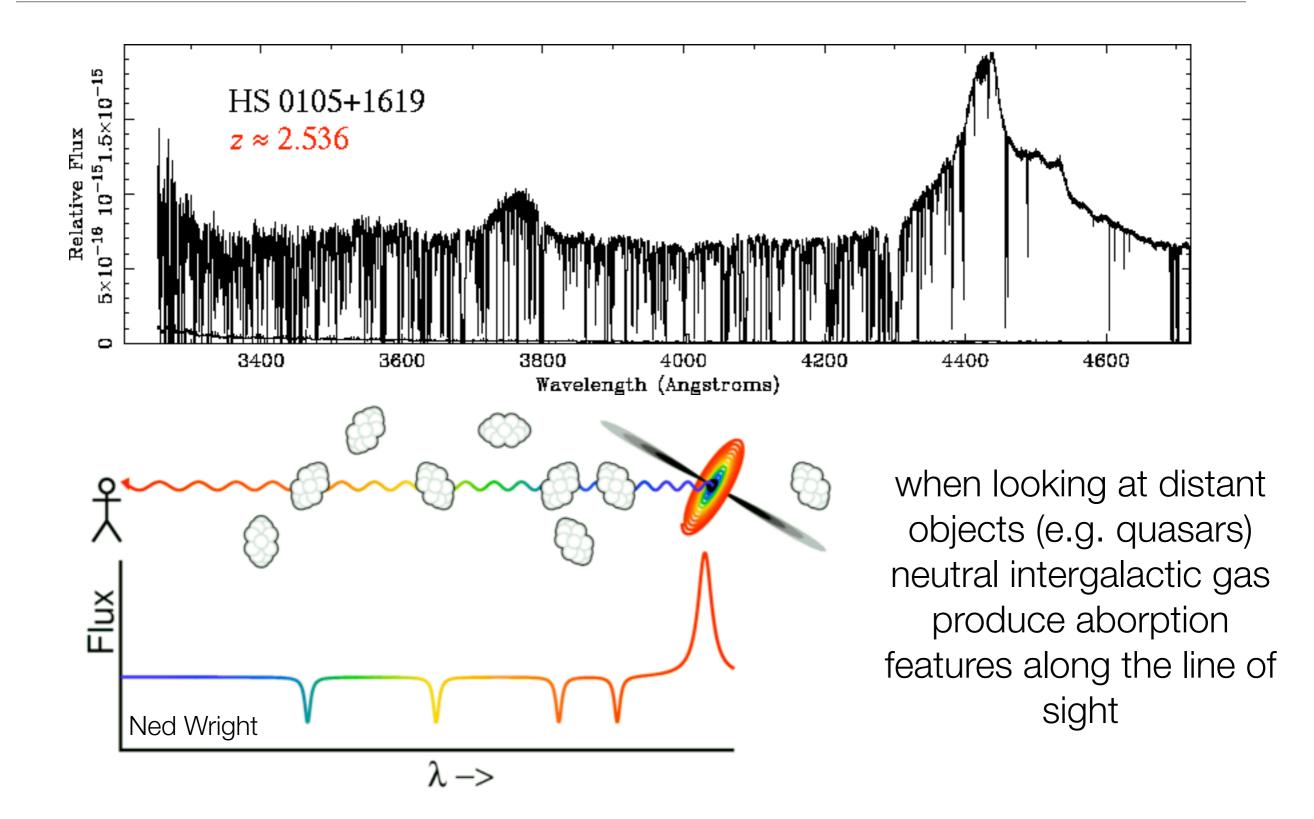
After recombination (t= 380 000 years), light & baryons are decoupled

Neutral Gas cools down and collapses because of gravity

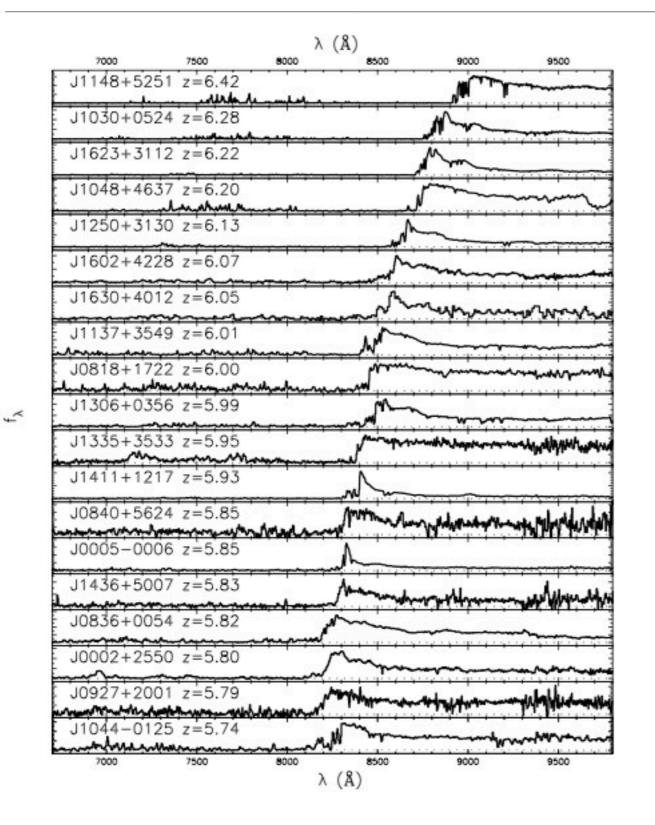
Roughly 200 Myrs after the BB, the density is large enough to convert hydrogen into stars

In between, no light is produced, the so-called Dark Ages

Lyman Alpha Forest



Gunn-Peterson Through

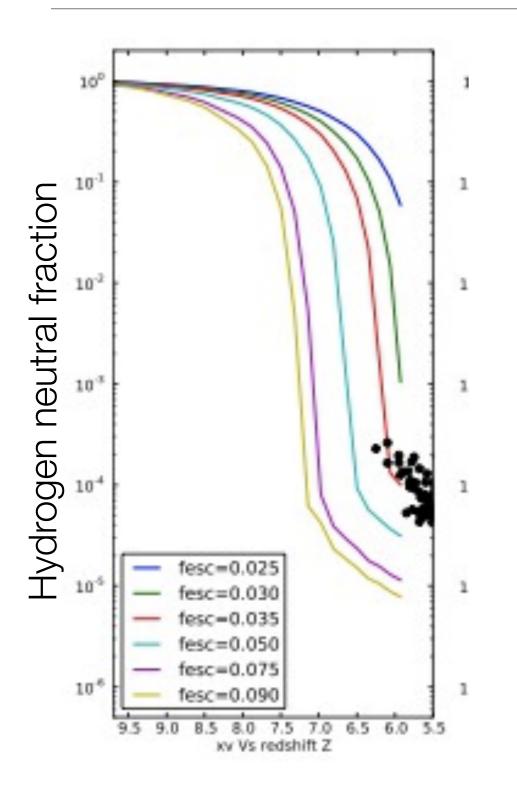


@ z>6 (1Gyr), spectras can be fully absorbed

no neutral windows with transmission

the Universe should be fully neutral at these times

Reionisation

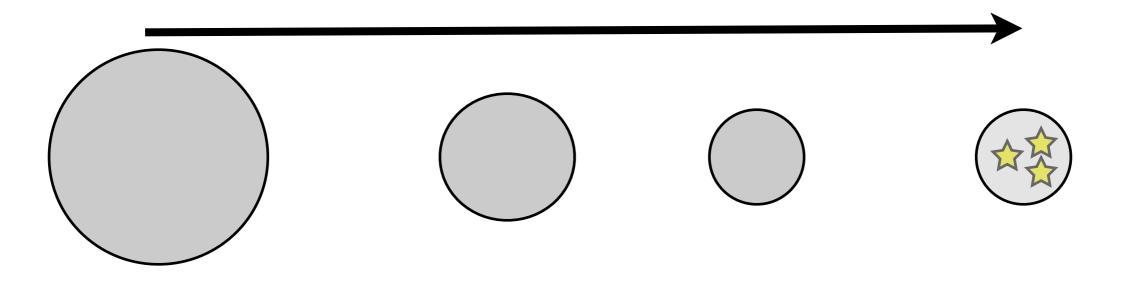


t=380 000 years: ionised -> neutral t=1Gyr : neutral ->ionised

The Universe becomes ionised again: The Reionisation

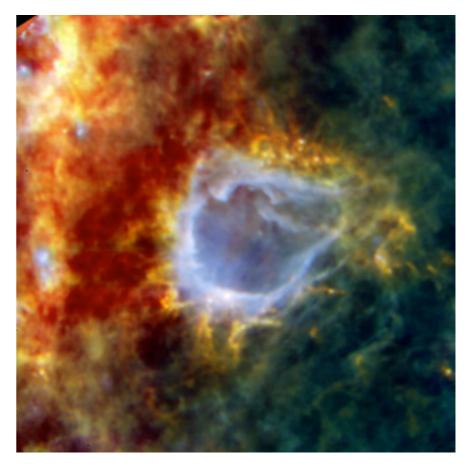
What happened ?

First Stars



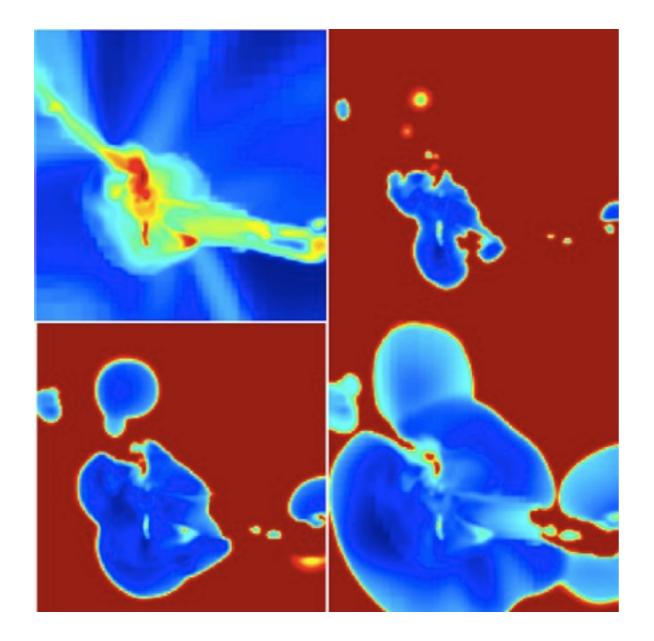
First Stars are expected to be metal-free, massive, and strong UV emitters

They appear in a gaseous environment



ESA/PACS/SPIRE/HOBYS Consortia

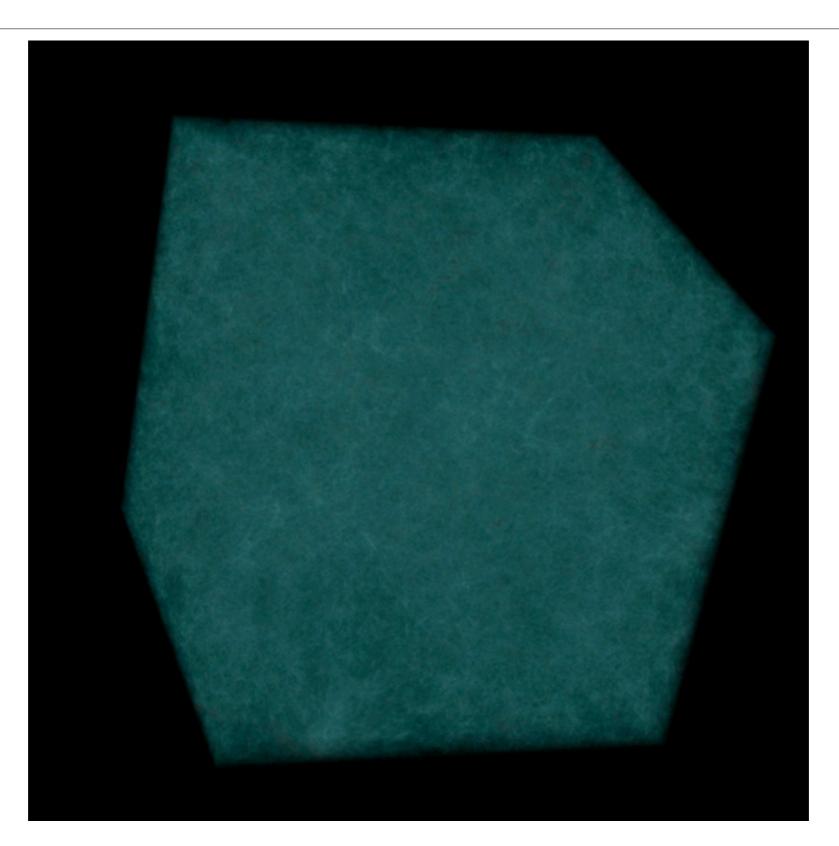
HII Regions



The new stars ionise their surroundings with their UV emission

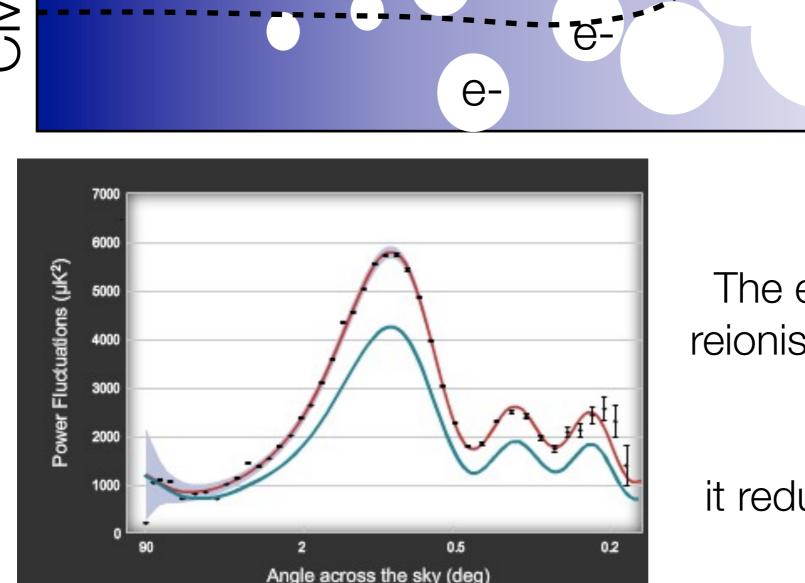
This process operates in the whole Universe

The last phase transition of the Universe



Reionisation & CMB





The electrons released by reionisation will scatter CMB photons

e-

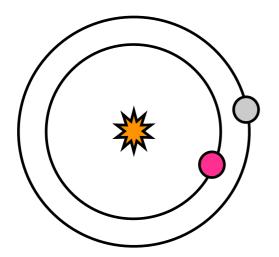
e-

it reduces the amplitude of fluctuations

Nasa/WMap

Matter (or the story of the first 8 Gyrs)

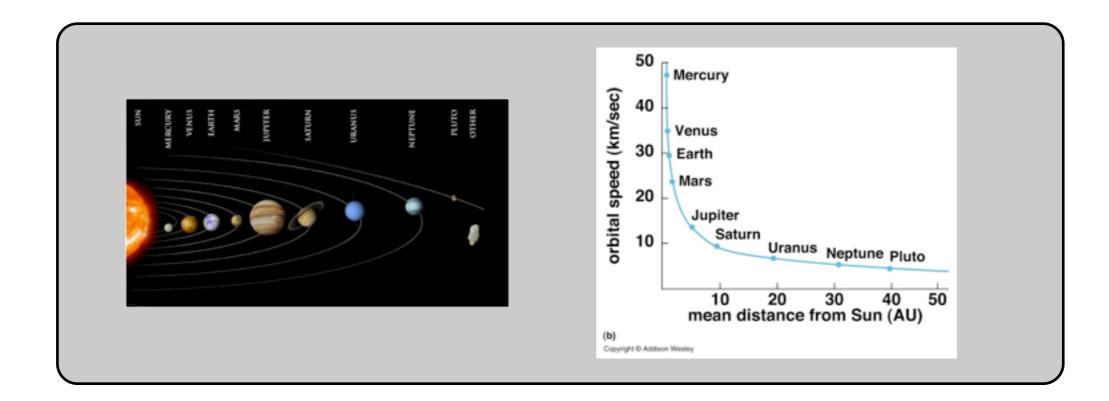
Mass & Rotation



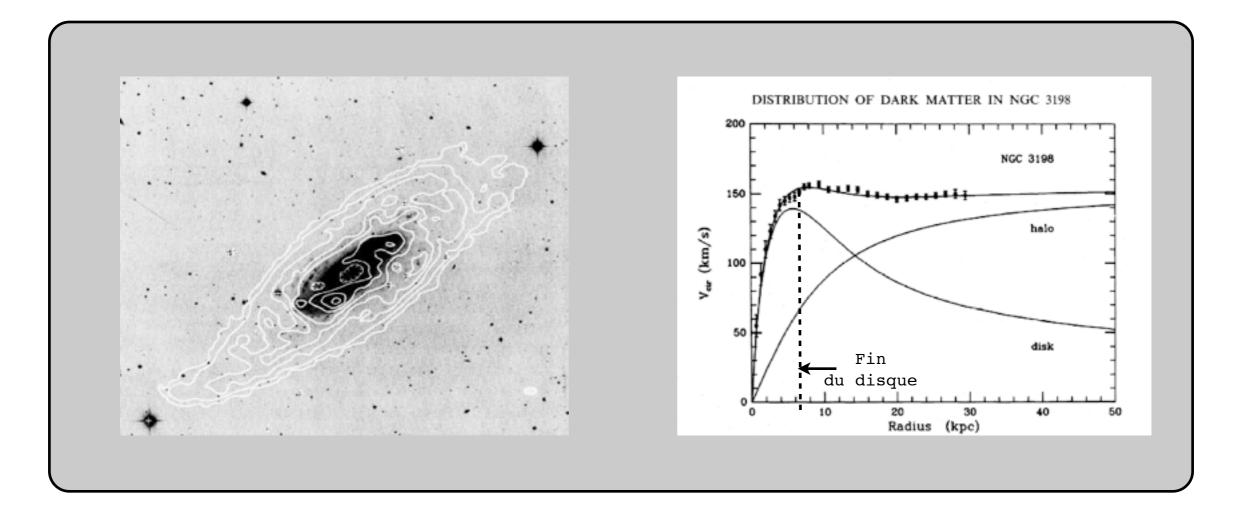
velocity of \bigcirc < velocity of \bigcirc

velocity depends on : - the radius of the orbit

- the mass enclosed in the orbit



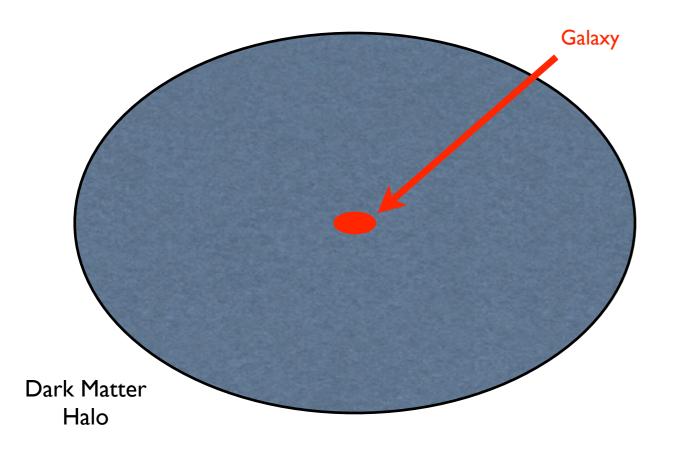
Rotation Curve of Galaxies



@ large distance, velocities are larger than expected

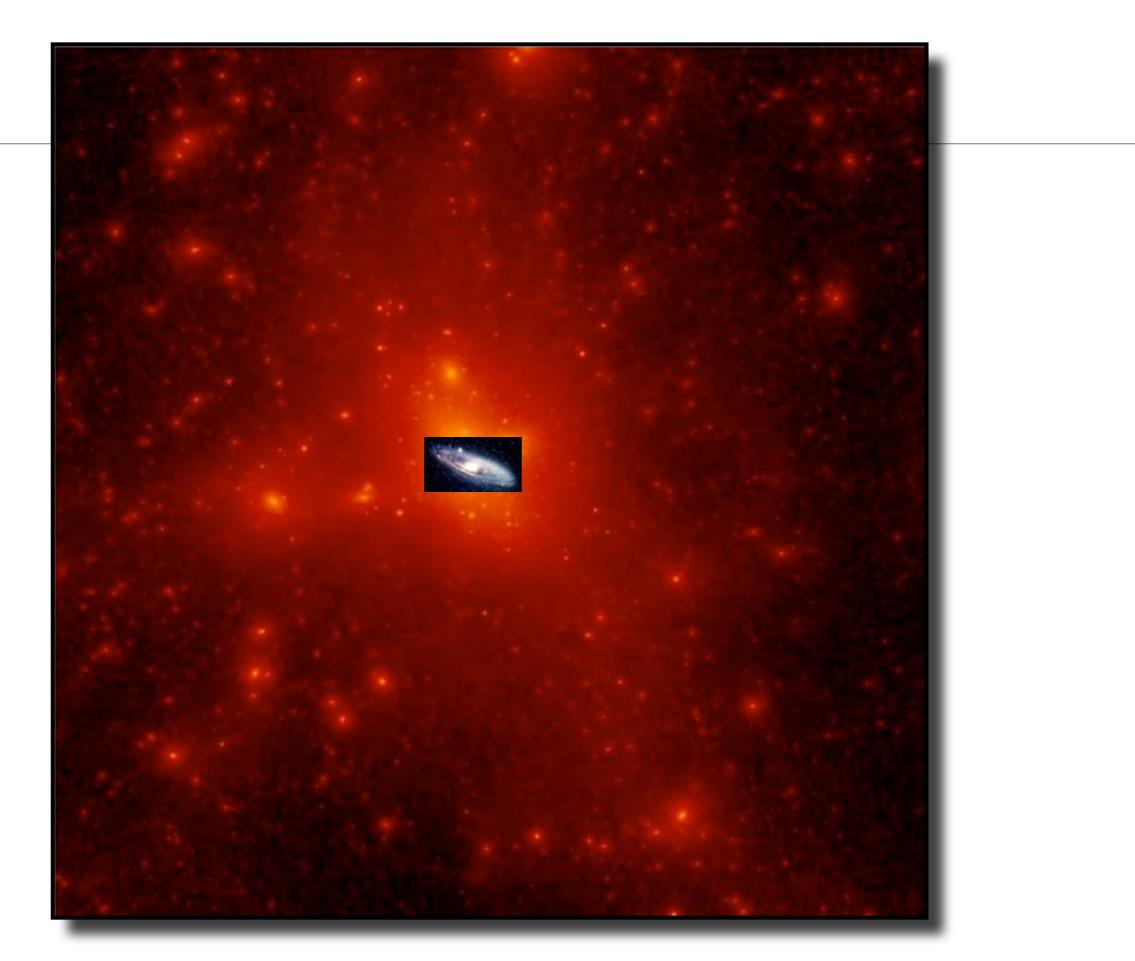
Dark Matter Haloes

Distant gas goes faster than expected. The amount of central mass is larger than observed

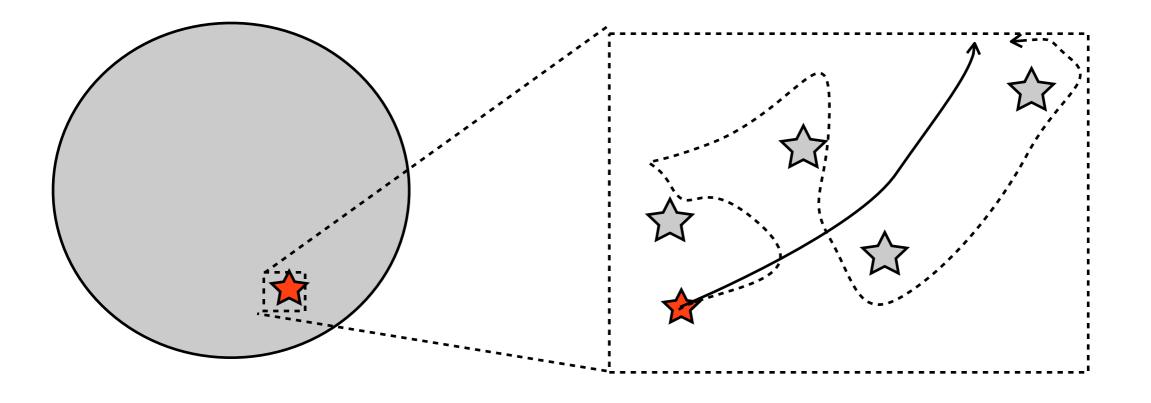


Dark Matter: gravitation only

~90% of the mass

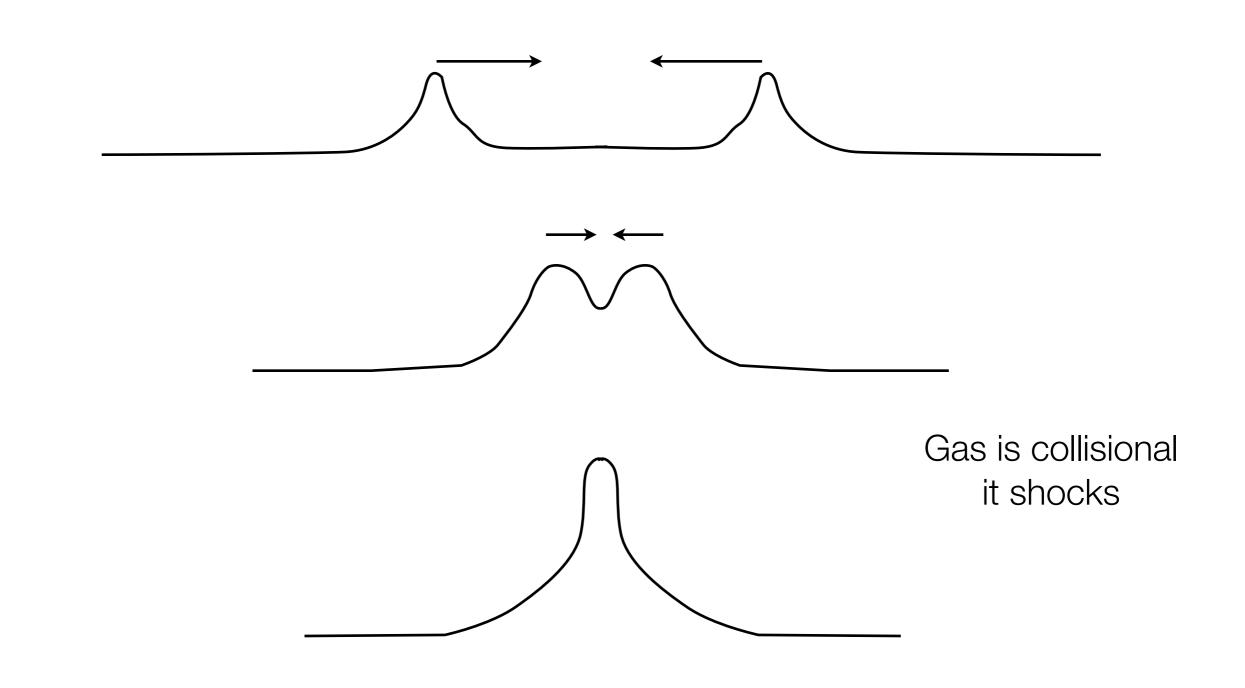


Collisionless Dark matter

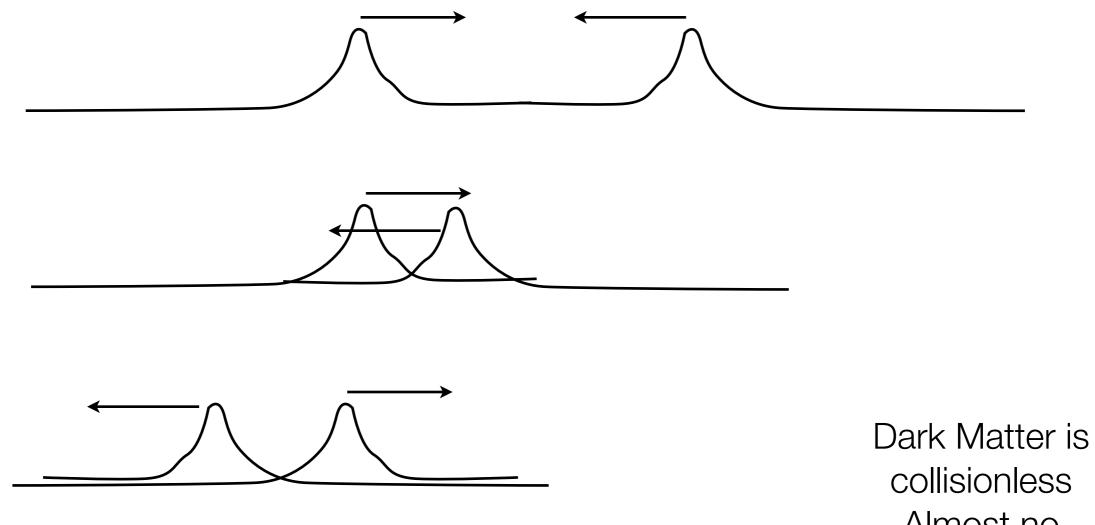


A dark matter particle is weakly influenced by its neighbors

Collisional Matter

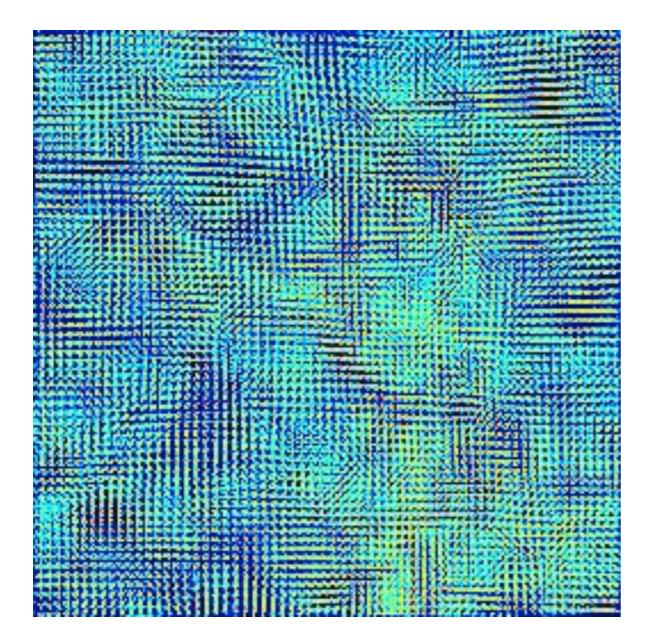


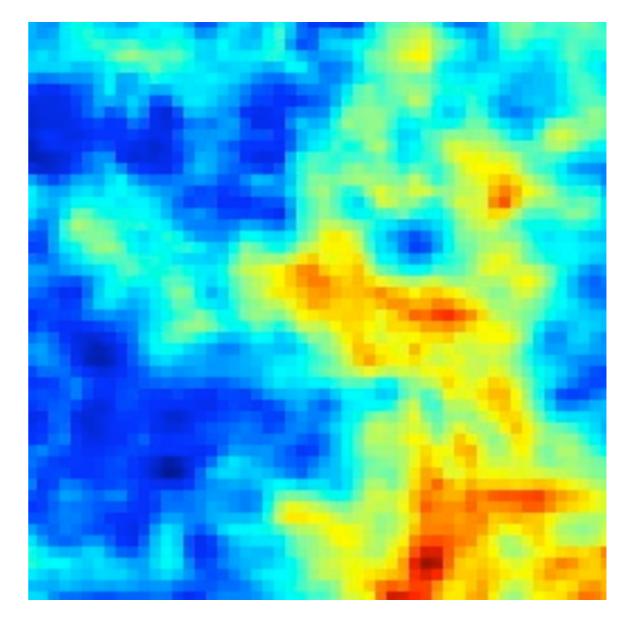
Collisionless Dark Matter



collisionless Almost no collisions

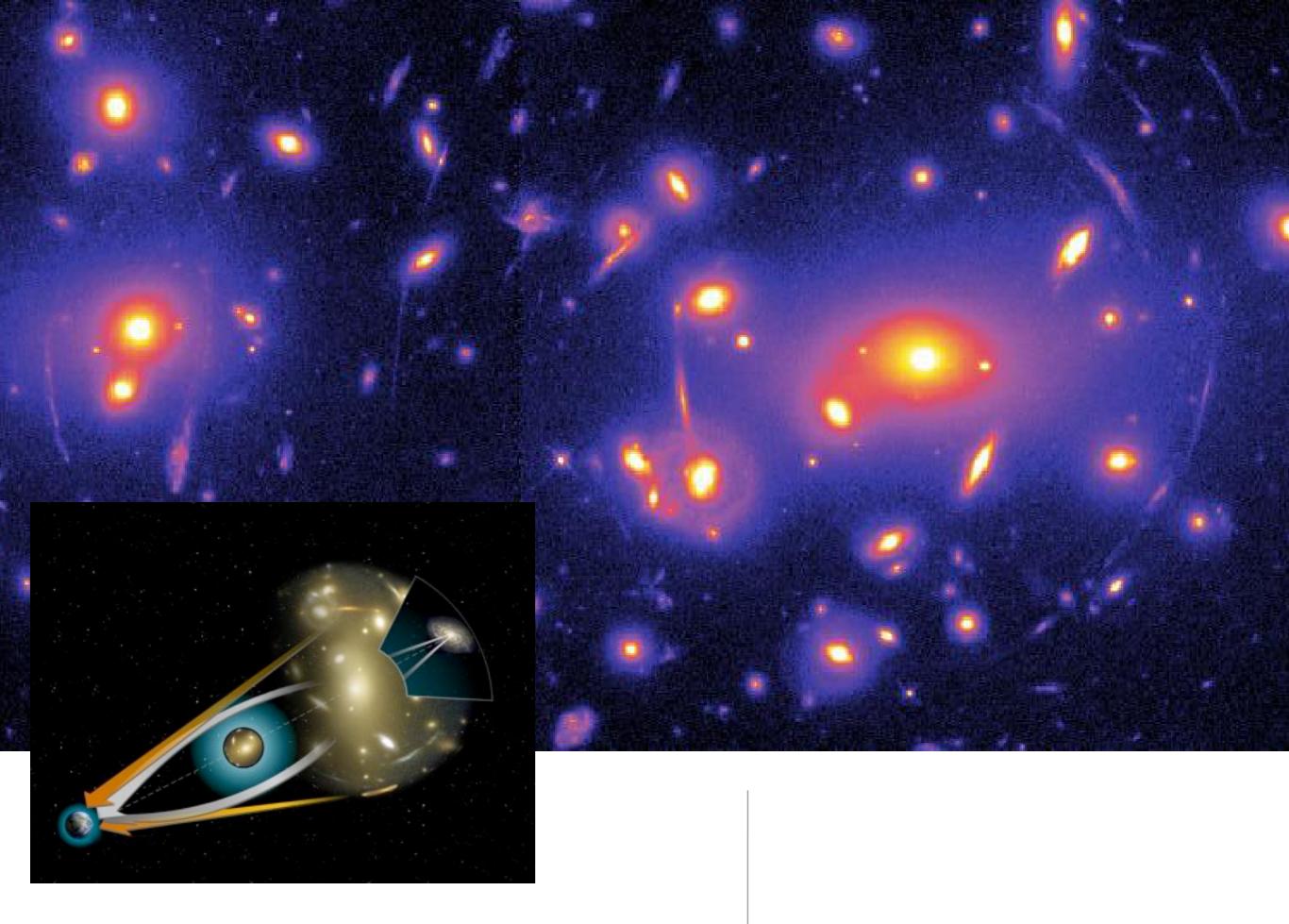




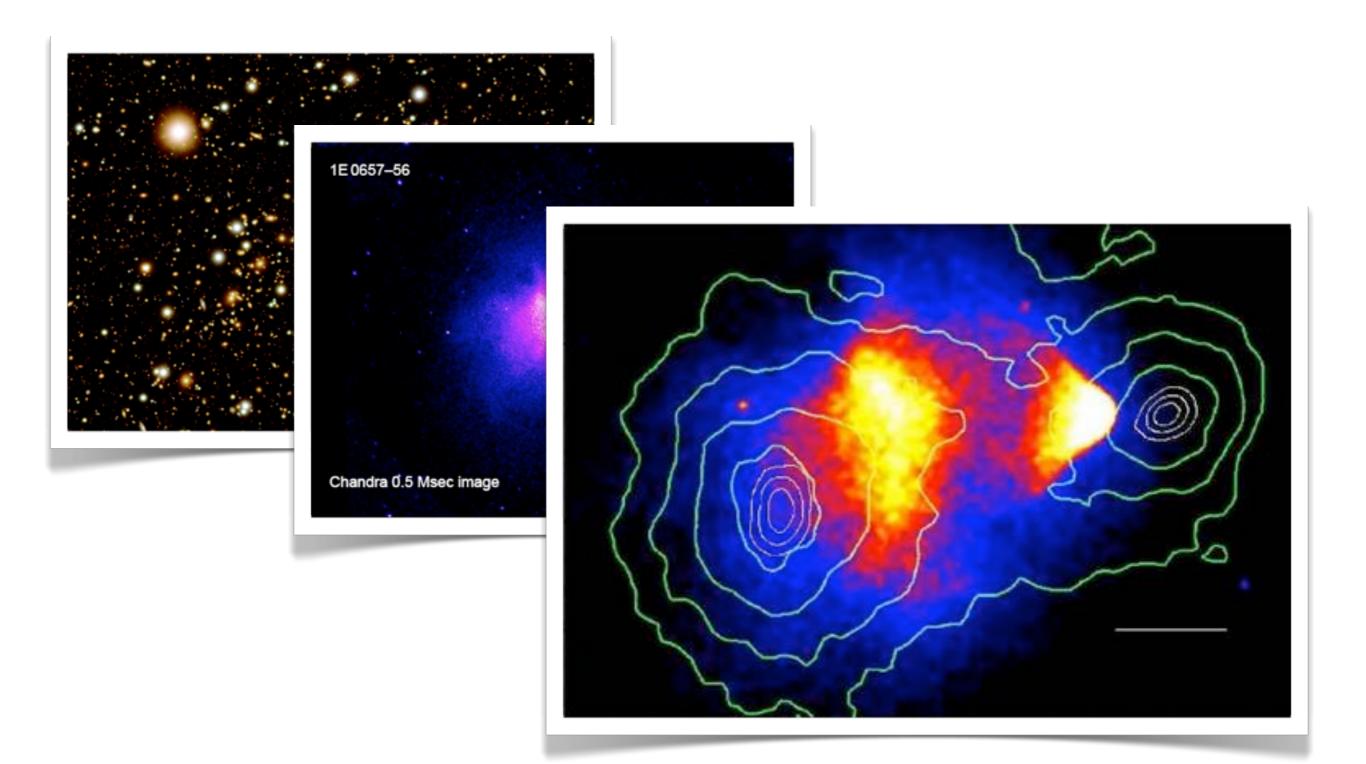


Dark Matter

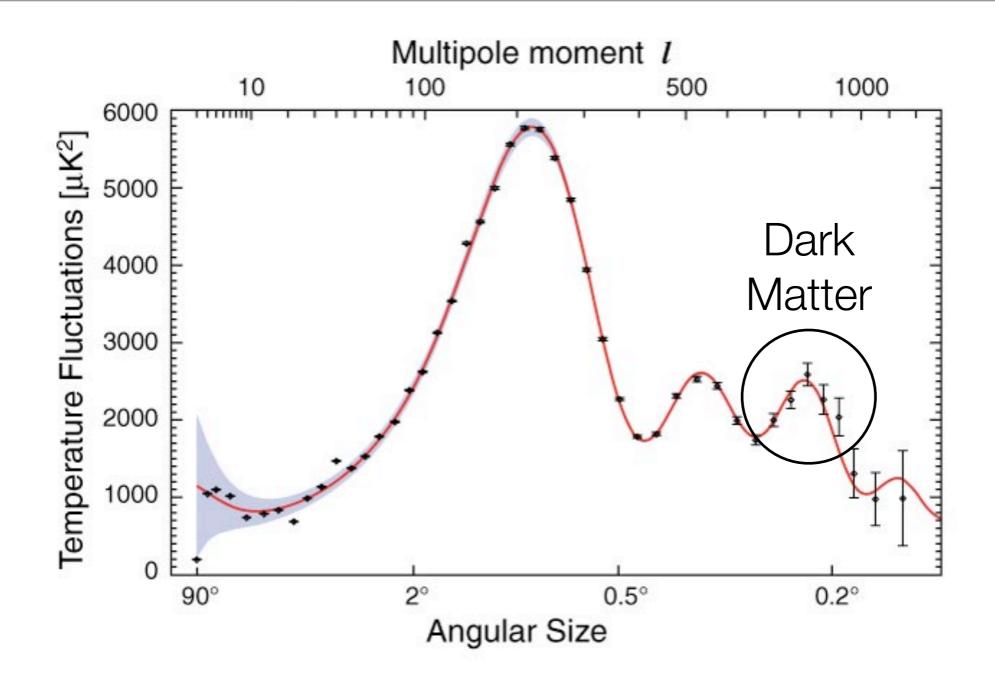




Invisible Matter



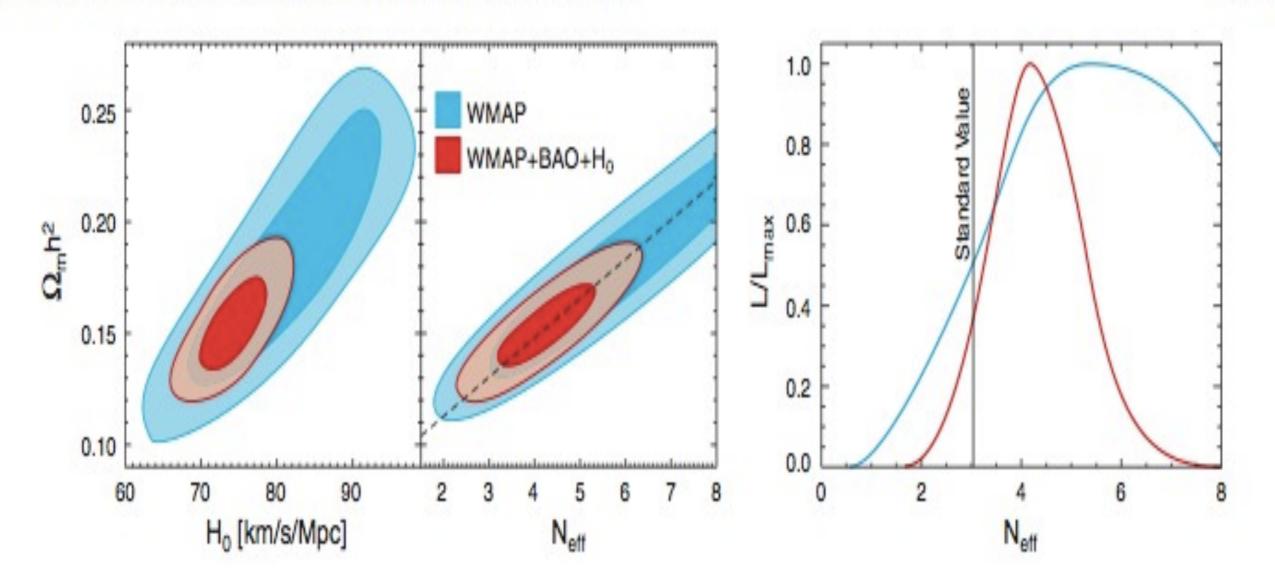
DM in CMB



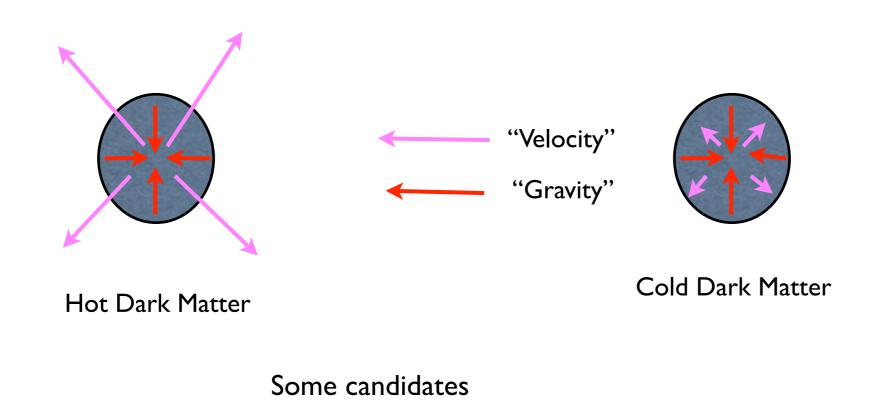
Number of relativistic species

OPHYSICAL JOURNAL SUPPLEMENT SERIES, 192:18 (47pp), 2011 February





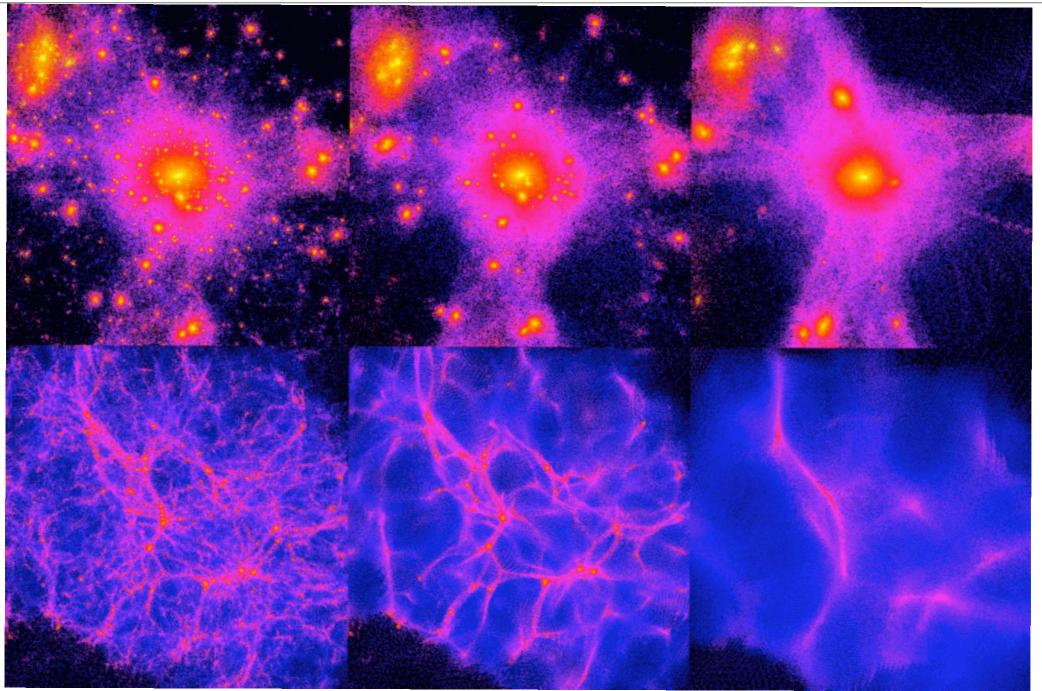
Nature of DM



-Neutrinos ? Sterile Neutrinos ?

- Cold Gas
- Supersymmetric particle
- Modified Gravity Beyond Einstein?

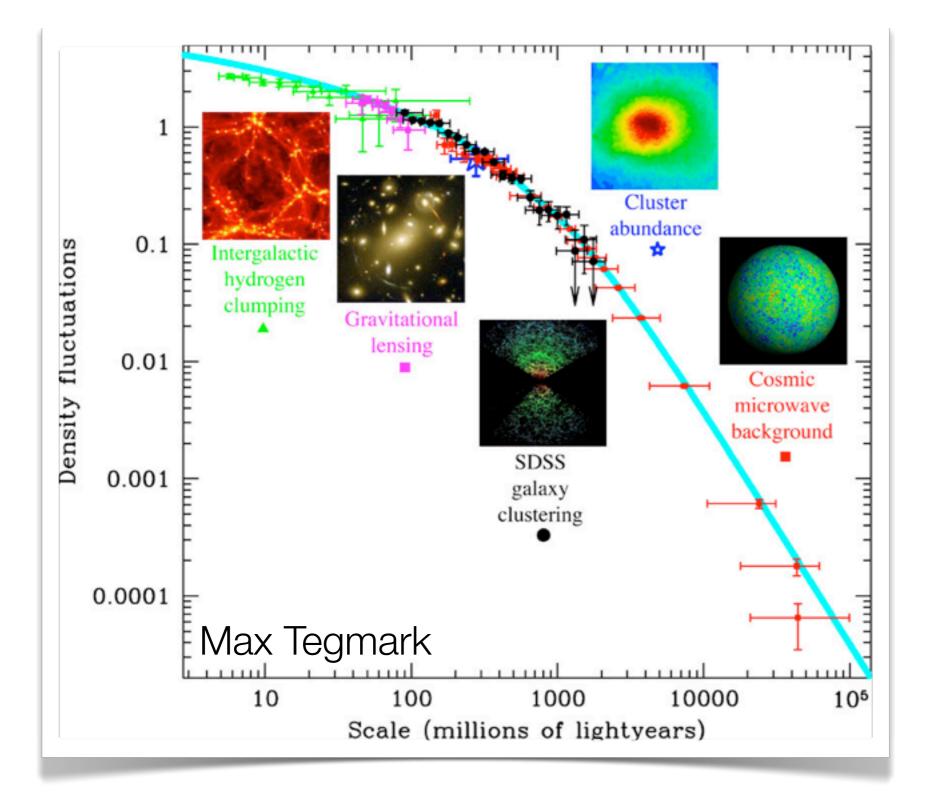
Warm Dark Matter Vs Cold Dark Matter



Ith Zurich

Small scales are suppressed in WDM

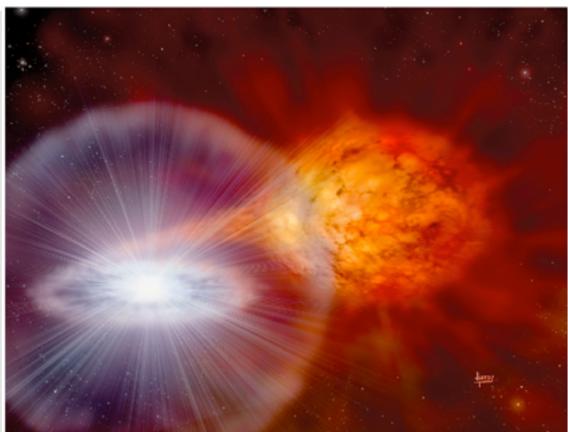
LSS and Dark Matter



Vacuum (or the story of the last 5 Gyrs)

SN IA

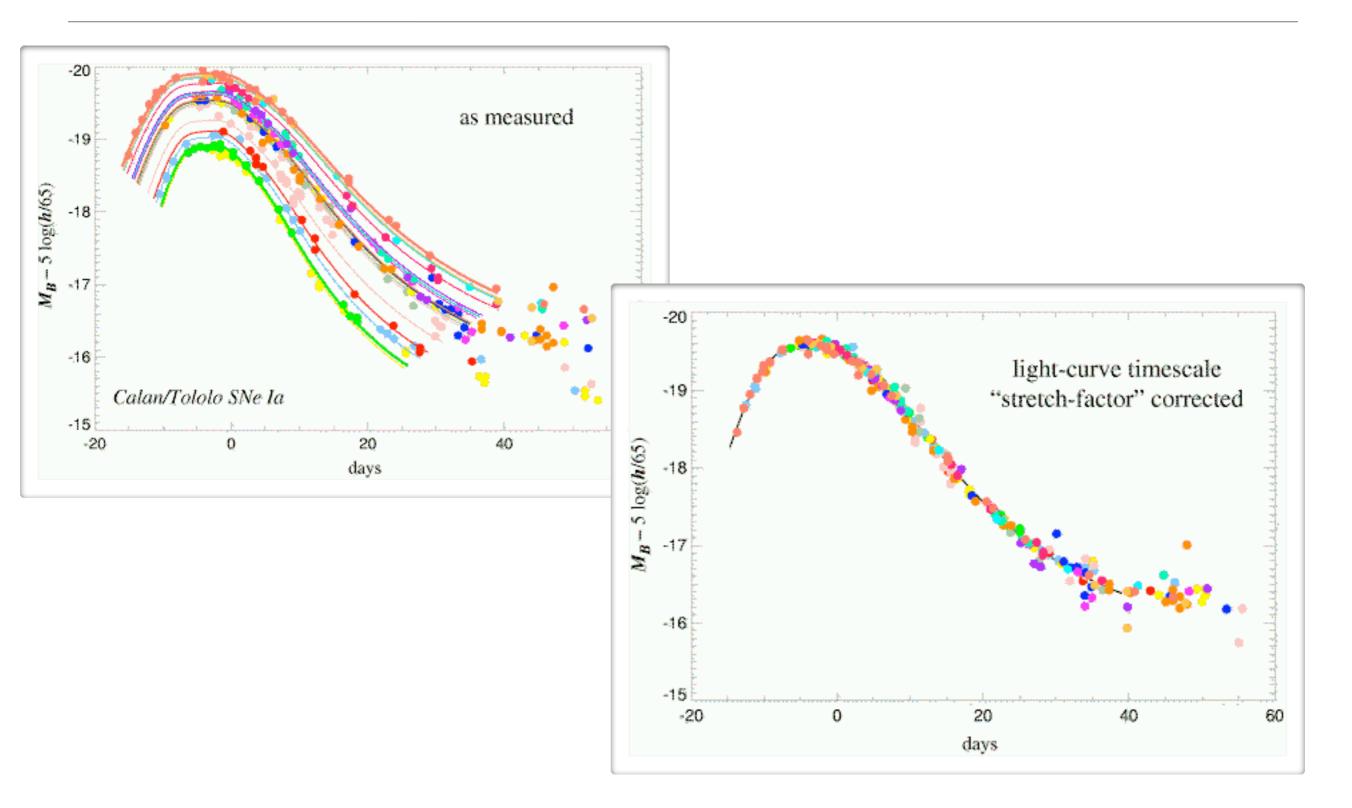


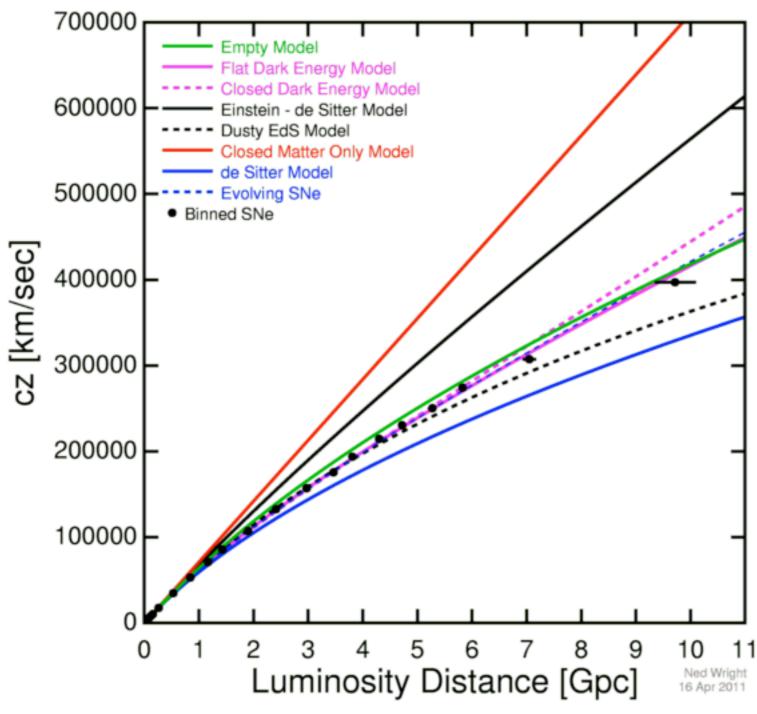


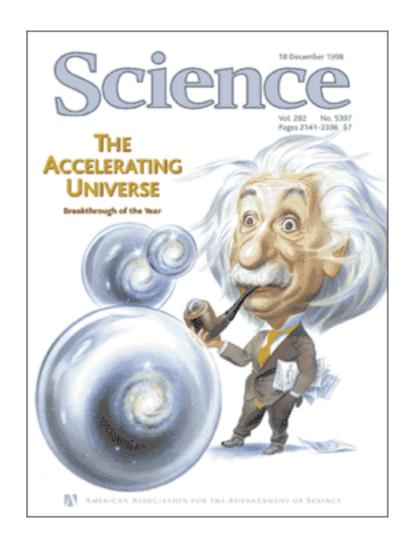
Artist's rendition of a white dwarf accumulating mass from a nearby companion star. This type of progenitor system would be considered singlydegenerate.

Image courtesy of David A. Hardy, © David A. Hardy/www.astroart.org.

SN Light Curves

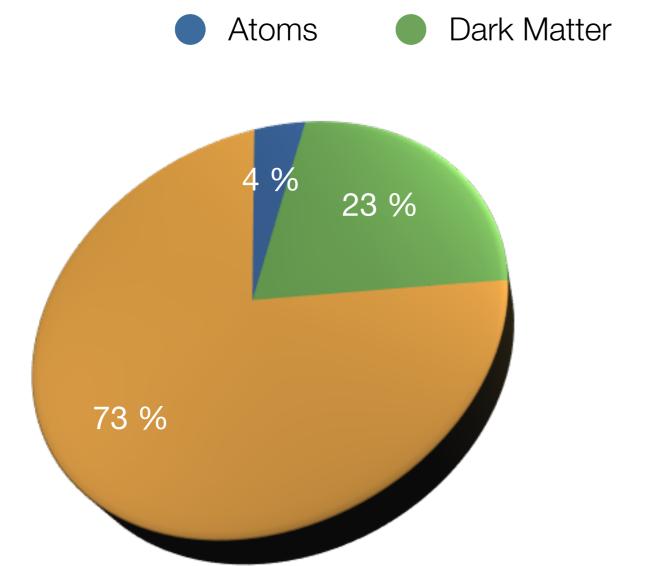






Ned Wright

Energetic content of the Universe, today



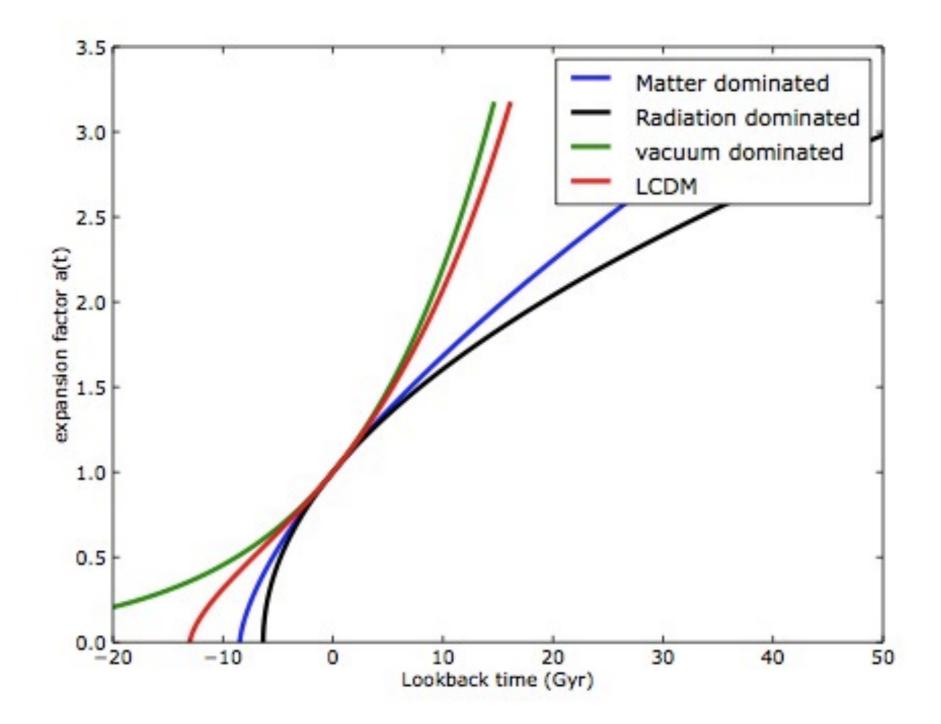
 $\Omega_{\Lambda} = 0.73$ $\Omega_{m} = 0.27$ $\Omega_{b} = 0.04$ $\Omega_{r} \sim 0.00001$

$$\Omega_m + \Omega_r + \Omega_\Lambda = 1$$

Vacuum Energy



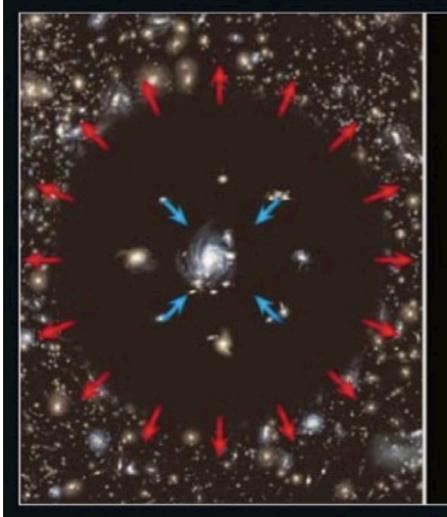
From energy to expansion



Big Rip

THE APOCALYPSE OF KNOWLEDGE

The accelerating cosmic expansion is beginning to undermine the three observational pillars of the big bang theory: the motion of galaxies away from one another, the cosmic microwave background radiation, and the relative quantities of light chemical elements such as hydrogen and helium.



BILLIONS OF YEARS LATER nearby galaxies have merged and distant ones have receded from view. The background radiation is undetectably dilute. Multiple generations of stars have contaminated the original chemical mix.

Supergalaxy

TODAY all three pillars are prominent. We see distant galaxies recede from us (*red arrows*) as nearby ones pull tighter (*blue*); background radiation suffuses space; and cosmic gas largely retains the chemical mix produced early in the big bang.

To Conclude

Cosmological Principle ?

The standard cosmological model assumes :

the Universe is isotropic and homogeneous (on cosmological scales)
Gravitation is described by General Relativity (GR)

In order to get rid of DM and DE one may (see B. Famaey talk):

- relax the assumption of isotropy or homogeneity
- assume that GR is only an effective theory

Otherwise:

ЛСЛ